

Latest Results on Searches for Dark Matter from IceCube

Matthias Danninger for the *IceCube Collaboration*

The Oskar Klein Centre for Cosmoparticle Physics, Stockholm University

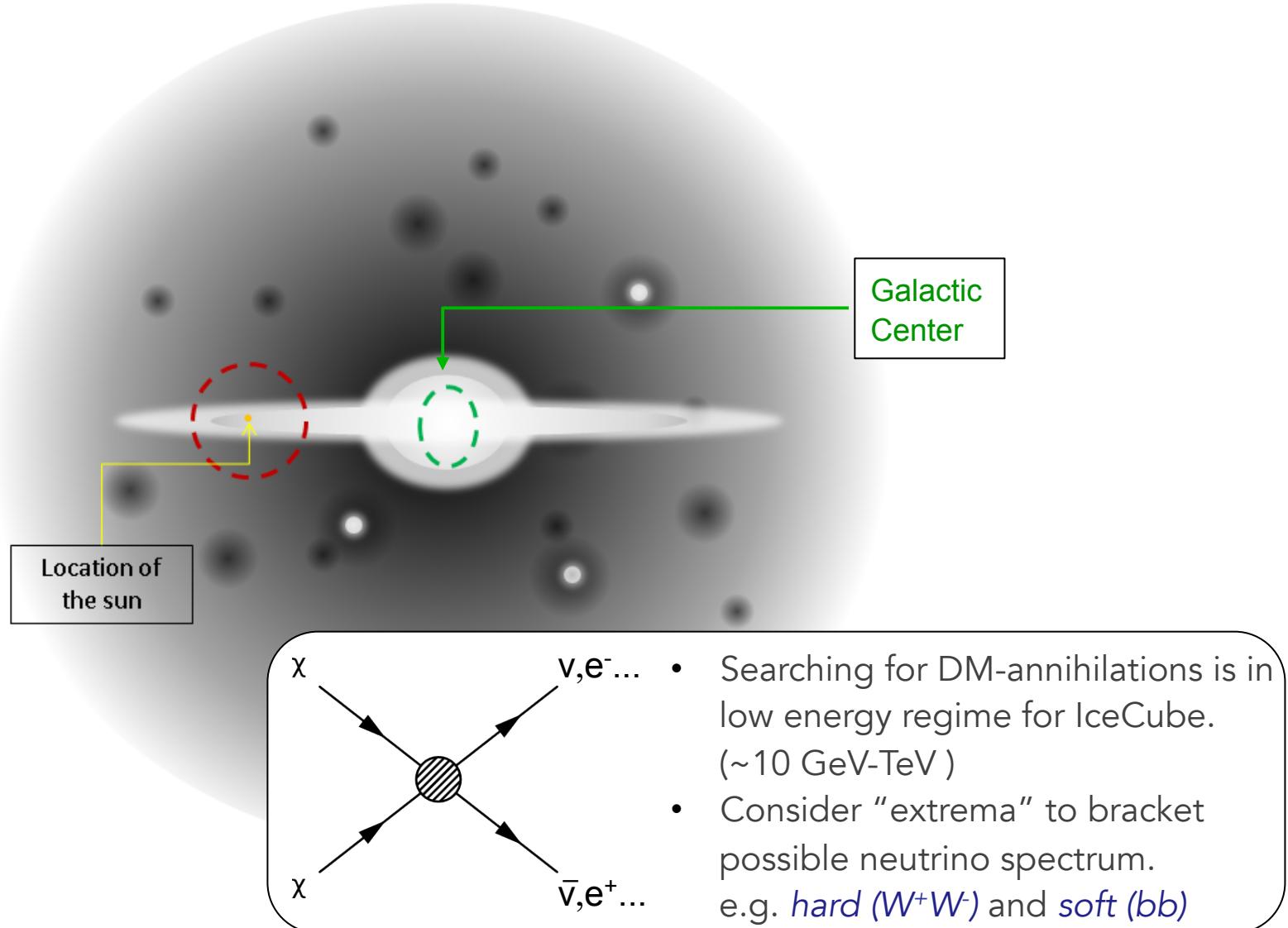
TAUP 2013 Topics in Astroparticle and Underground Physics
Asilomar, September 8-13, 2013



Indirect Search with IceCube (Overview)



(Image: M.Strassler)



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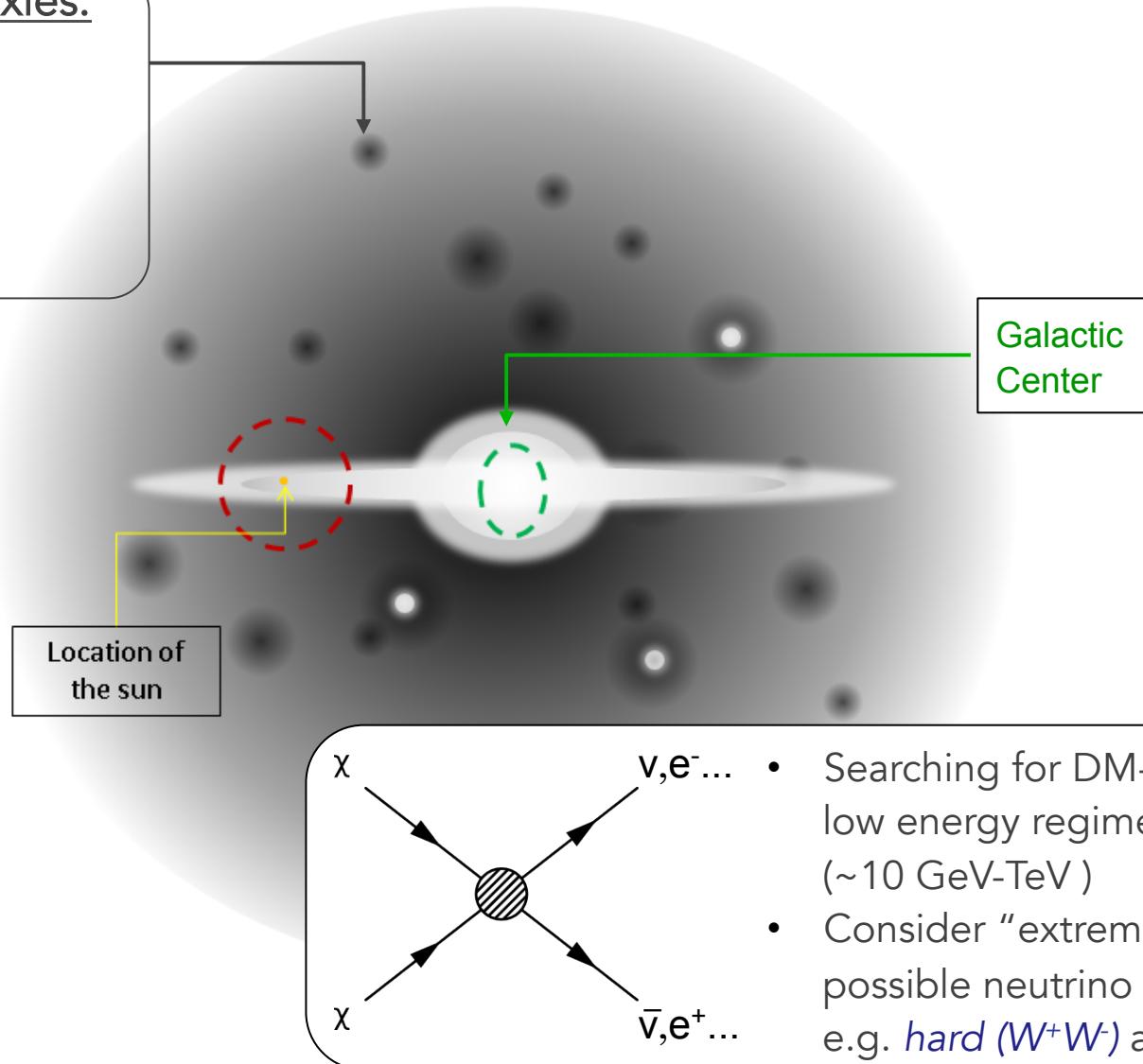
Dwarf spheroidal Galaxies:

→ IceCube-59 limits

Clusters of Galaxies:

→ IceCube-59 limits

(arXiv:1210.3557 2012)



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Galactic Halo:

→ IceCube-22 limits

(PRD 84 (2011) 022004)

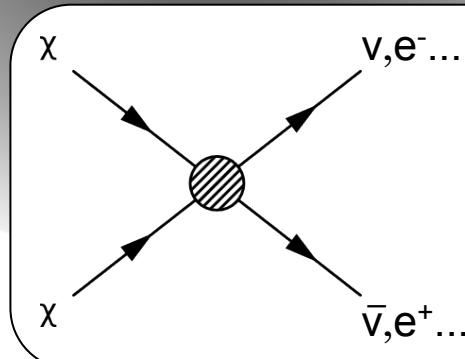
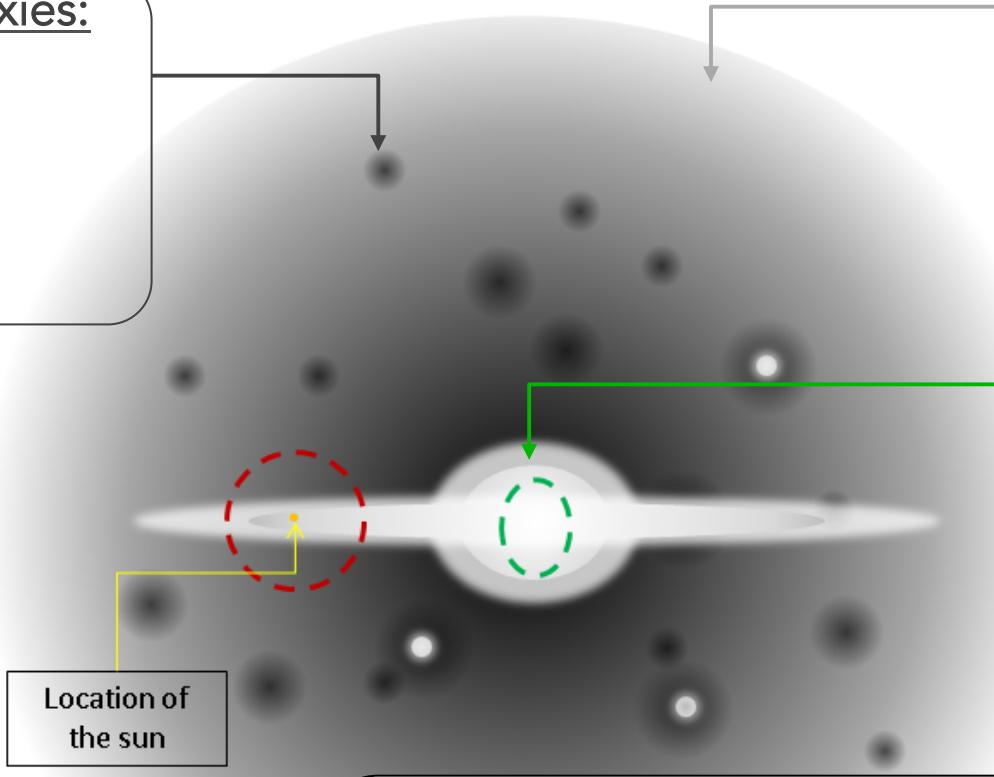
→ IceCube-79 limits

Galactic Center:

→ IceCube-40 limits

(arXiv:1210.3557 2012)

→ IceCube-79 sens



- Searching for DM-annihilations is in low energy regime for IceCube. (~ 10 GeV-TeV)
- Consider “extrema” to bracket possible neutrino spectrum.
e.g. *hard* (W^+W^-) and *soft* (bb)

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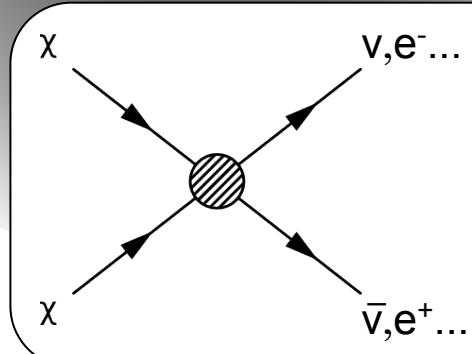
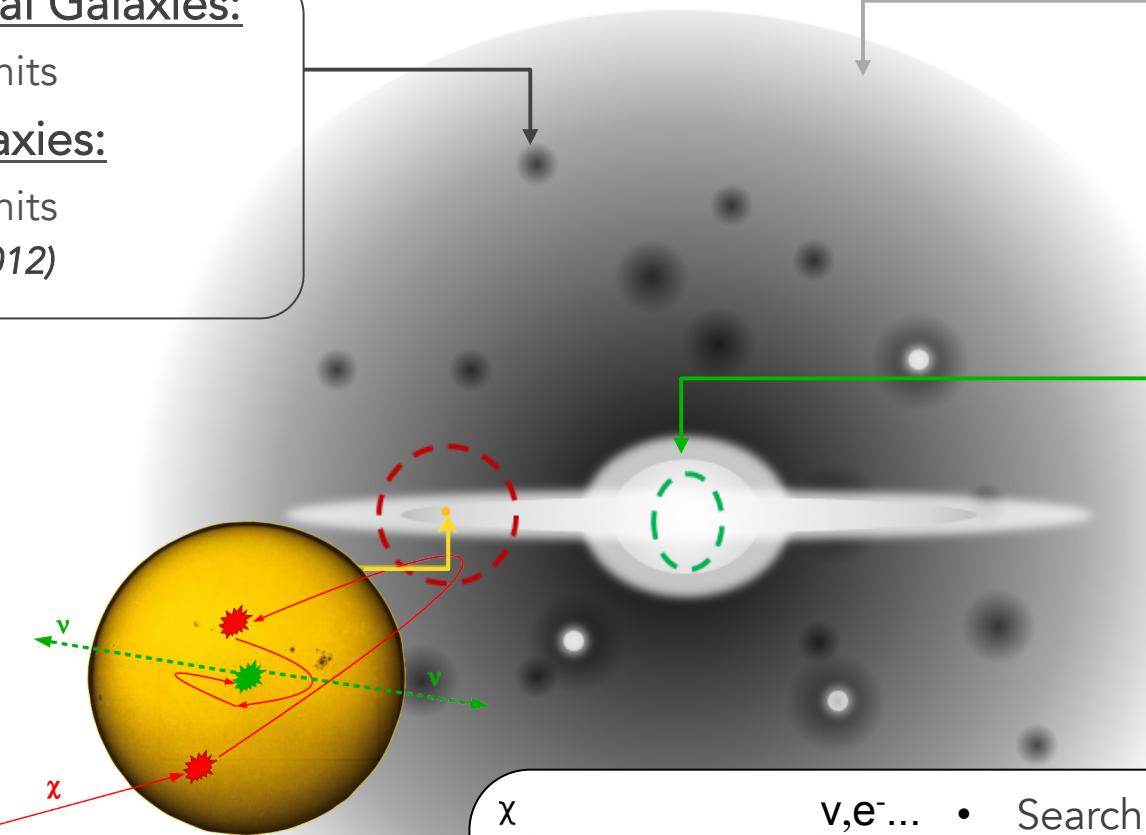
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Local sources (Sun & Earth):

→ IceCube-79 limits

(PRL 110 (2013) 131302)

→ Specific models & Global fits
(JCAP 11 (2012) 057)



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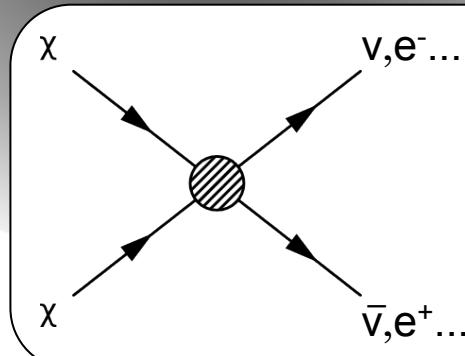
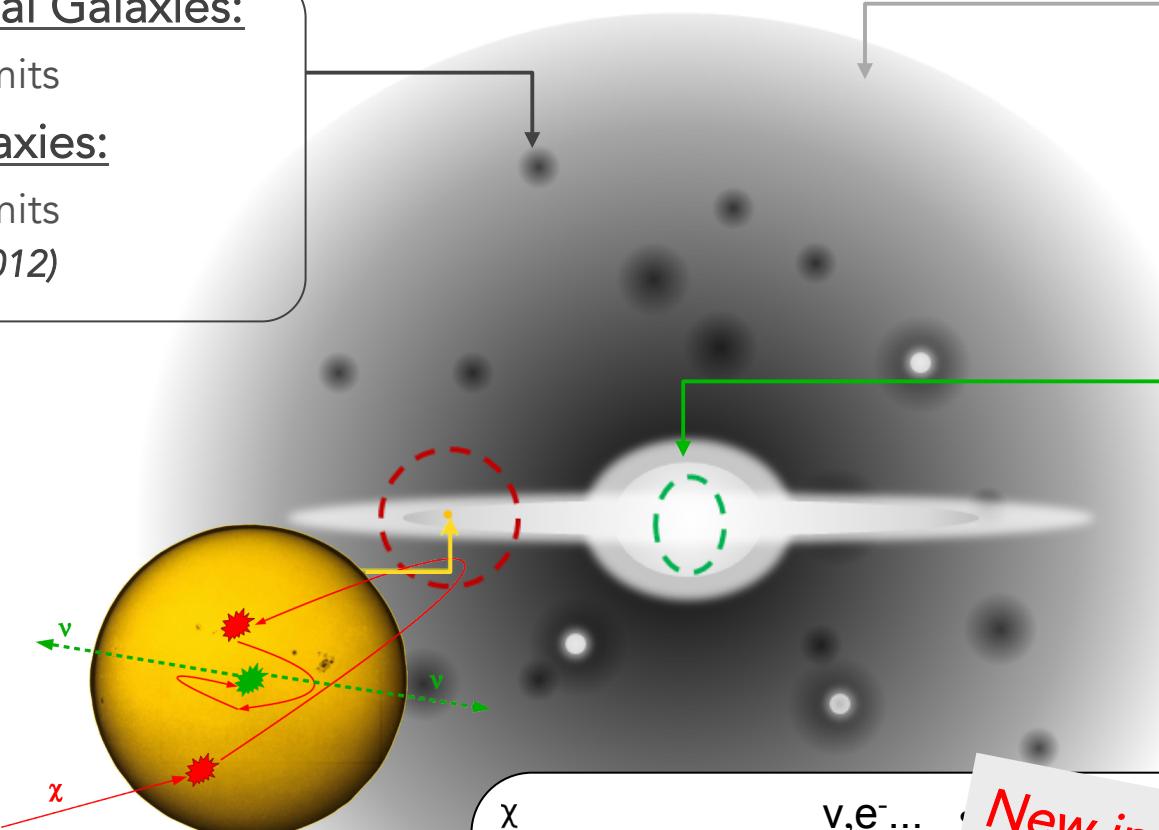
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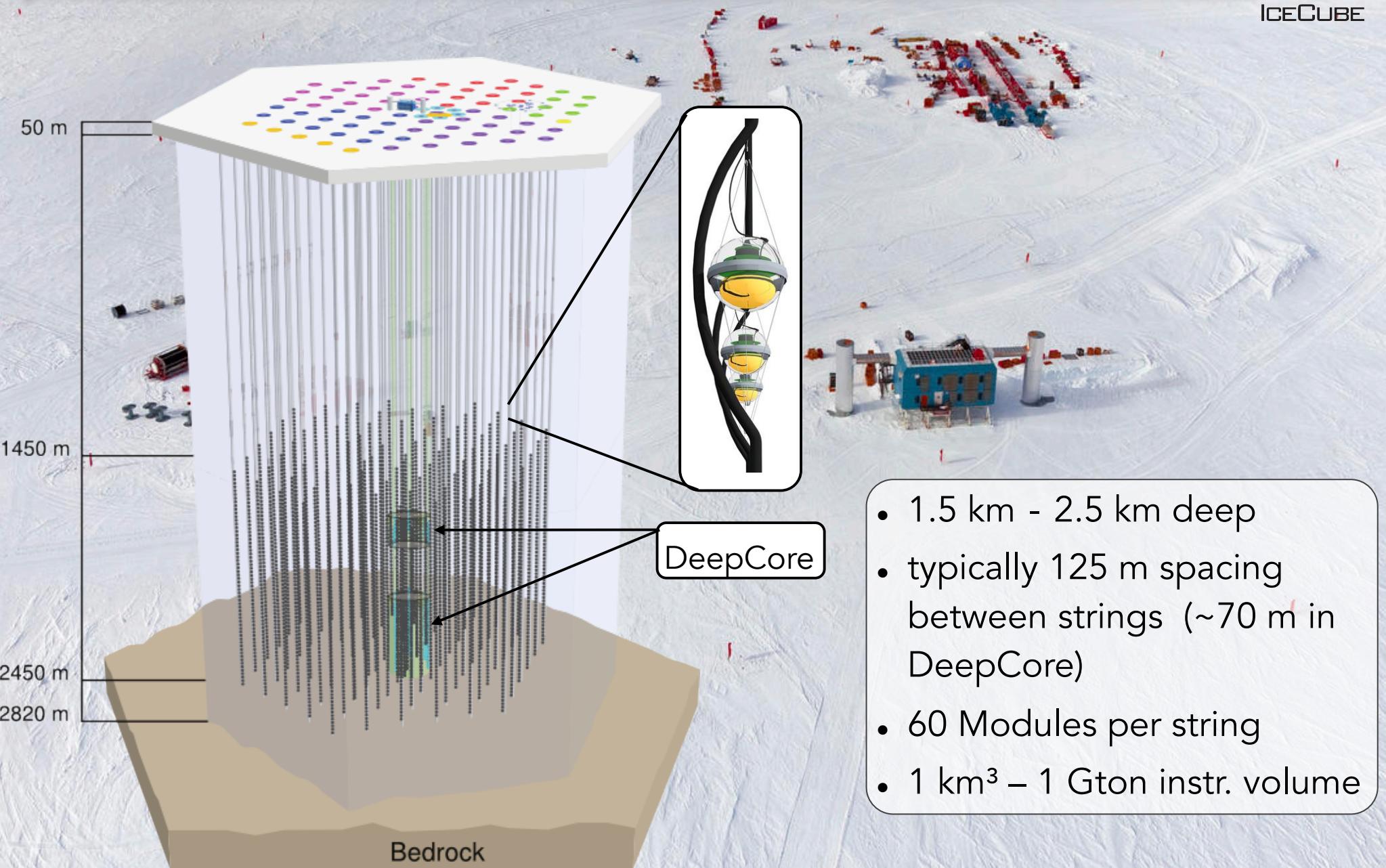
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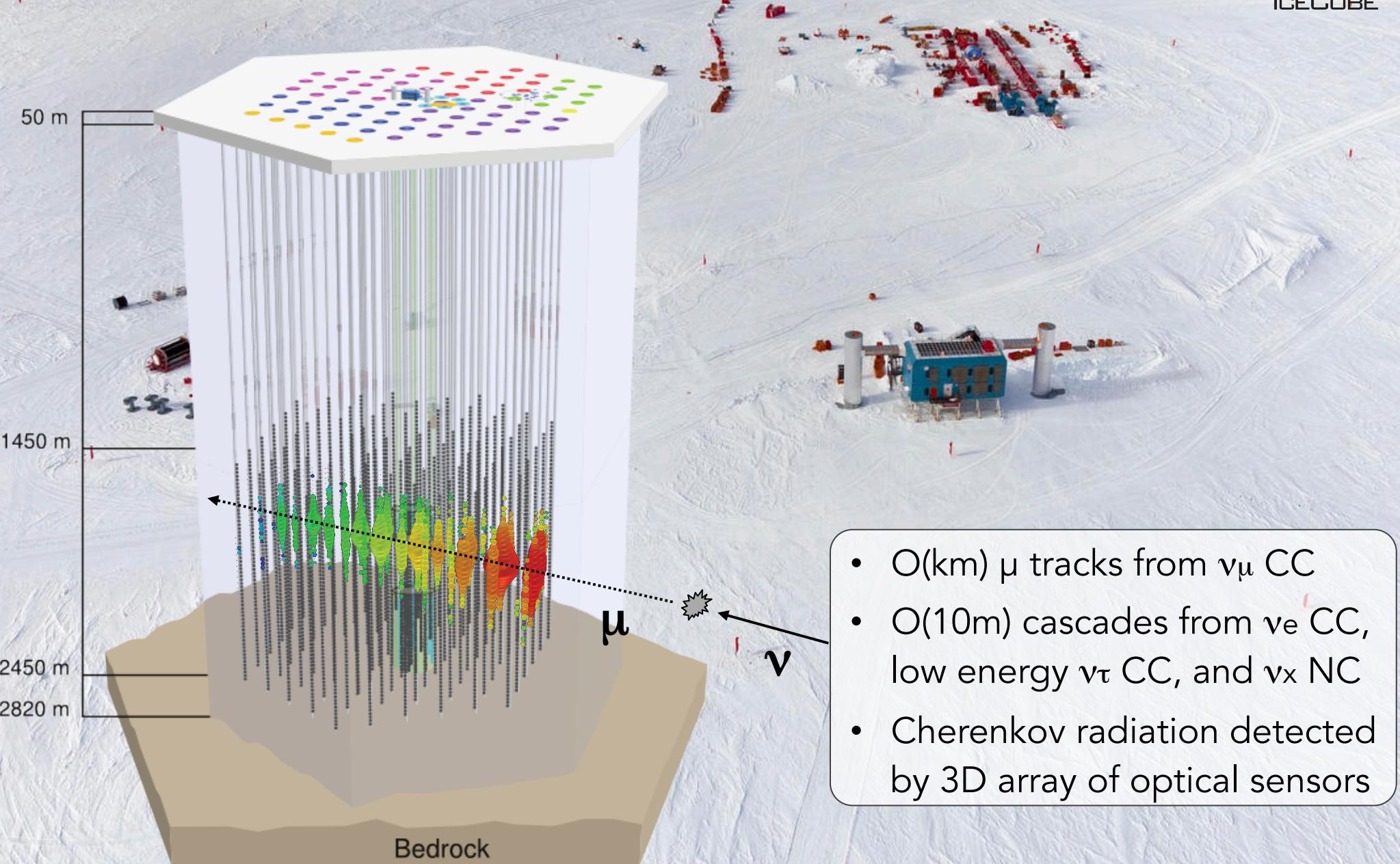
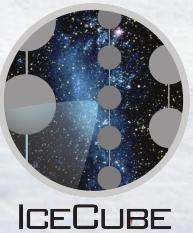
- New interest in PeV range
- DM-annihilations is in low energy (~10 GeV-TeV)
 - Consider "extrema" to bracket possible neutrino spectrum.
e.g. *hard* (W^+W^-) and *soft* (bb)

The IceCube detector



- 1.5 km - 2.5 km deep
- typically 125 m spacing between strings (~70 m in DeepCore)
- 60 Modules per string
- 1 km³ – 1 Gton instr. volume

The IceCube detector



Solar Dark Matter Search with IceCube



- All processes depend on WIMP mass
- Annihilation channel (branching ratios)
- Annihilation cross-section
- Capture (scattering)
→ Scattering cross-sections (SI & SD)

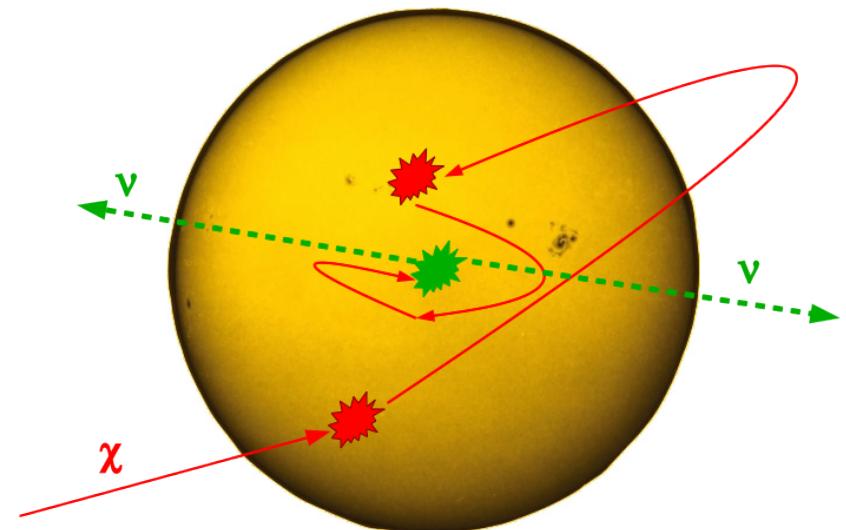
Details about Capture Process (e.g.):

Press & Spergel '85

Gould '88

Peter 2008

Sivertsson, Edsjö, PRD85 (2012) 123514



$$\frac{dN}{dt} = C_{capt.} - C_{ann.}$$

$$C_{ann.} = C_{capt.} \rightarrow \sigma_{total}$$

Proposed by:

Silk, Olive & Srednicki '85

Gaisser, Steigman & Tilav '86

Freese '86

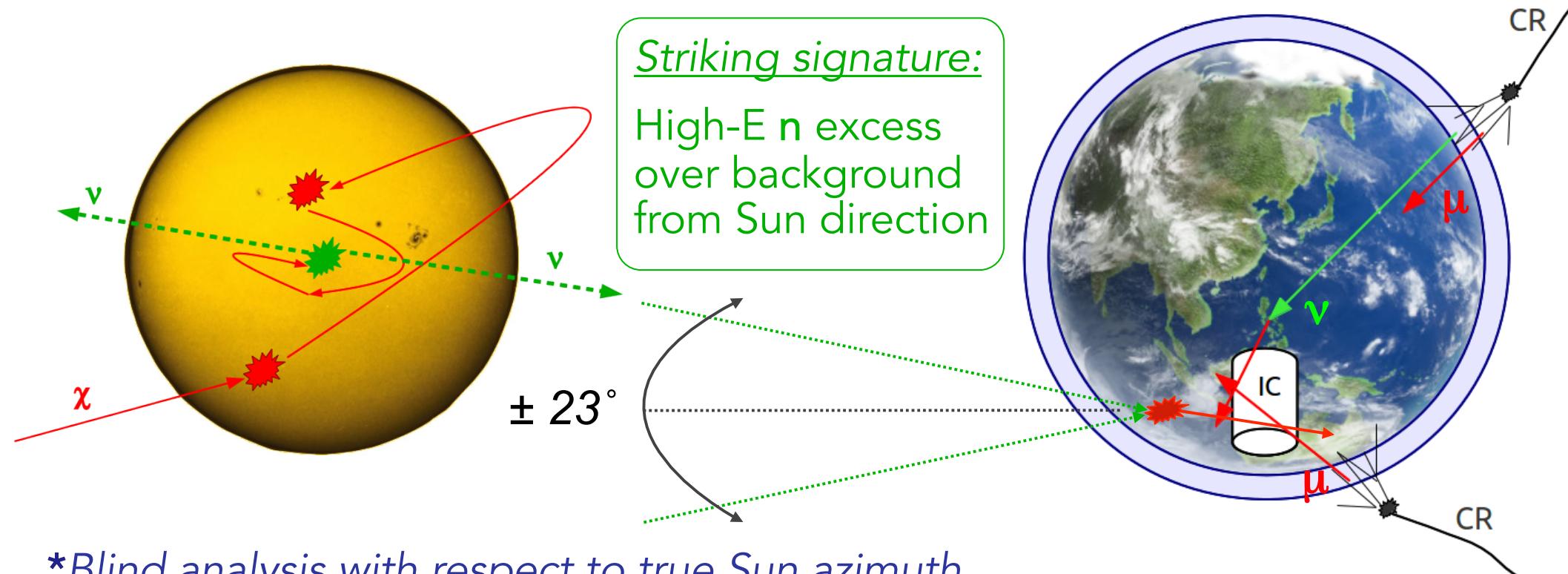
Krauss, Srednicki & Wilczek '86

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main analysis backgrounds:
 $\text{atm.}\mu \sim \mathcal{O}(10^8)$ triggering events/day
 $\text{atm.}\nu \sim \mathcal{O}(10^3)$ triggering events/day



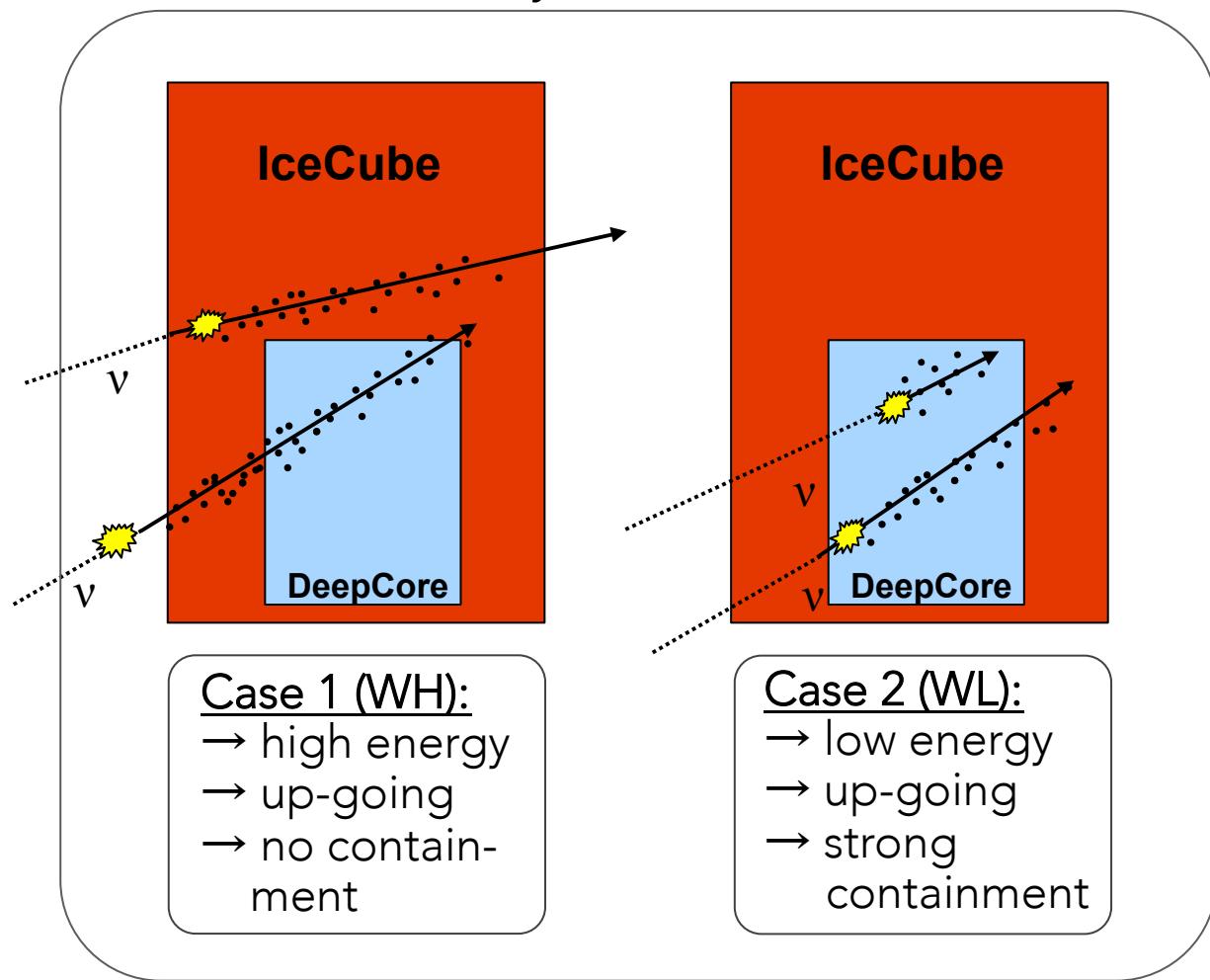
*Blind analysis with respect to true Sun azimuth

IceCube-79 string analysis details

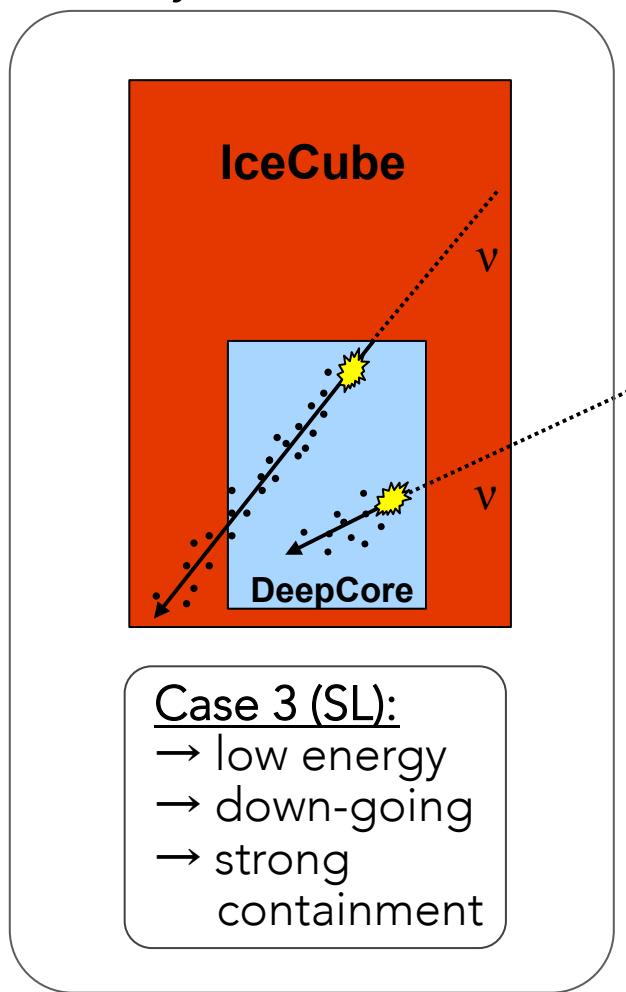


- Analysis for the whole year! Used 317 days livetime
- With DeepCore, analysis reaches neutrino energies of 10-20GeV

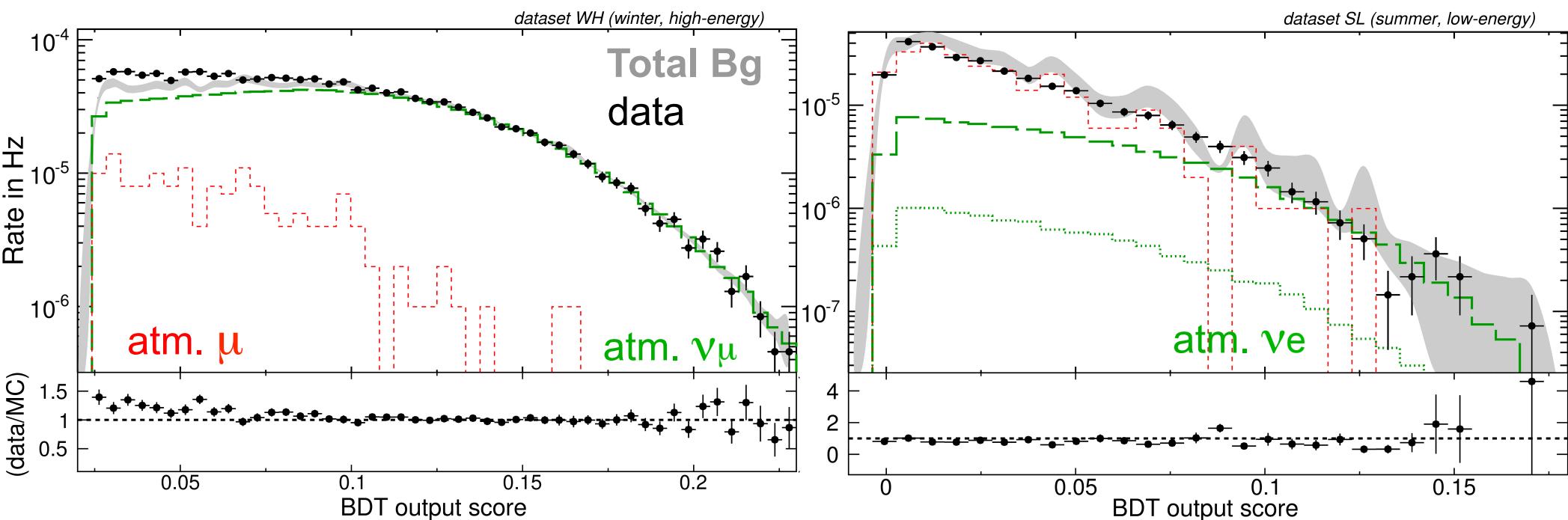
151 days austral winter



166 days austral summer



Multivariate analysis step (final cut applied)

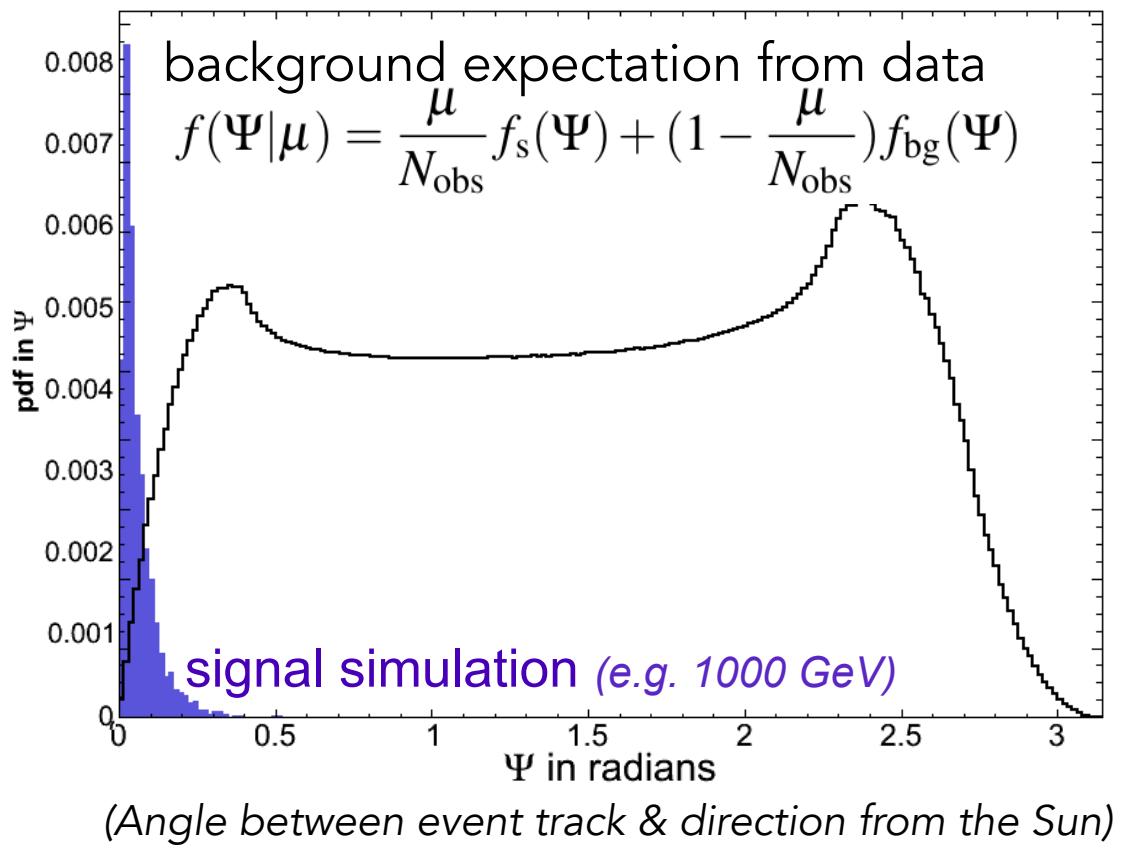


- 1 separate BDT for each event selection
- Training on off-source exp. data + separate signal simulation
- Optimized final cut on BDT-output: run llh-analysis for various BDT cuts; determine cut value with best sensitivity

Maximum likelihood-analysis



The observed angle to the Sun is fitted with *signal* and *background* pdf:s



Maximum likelihood-analysis

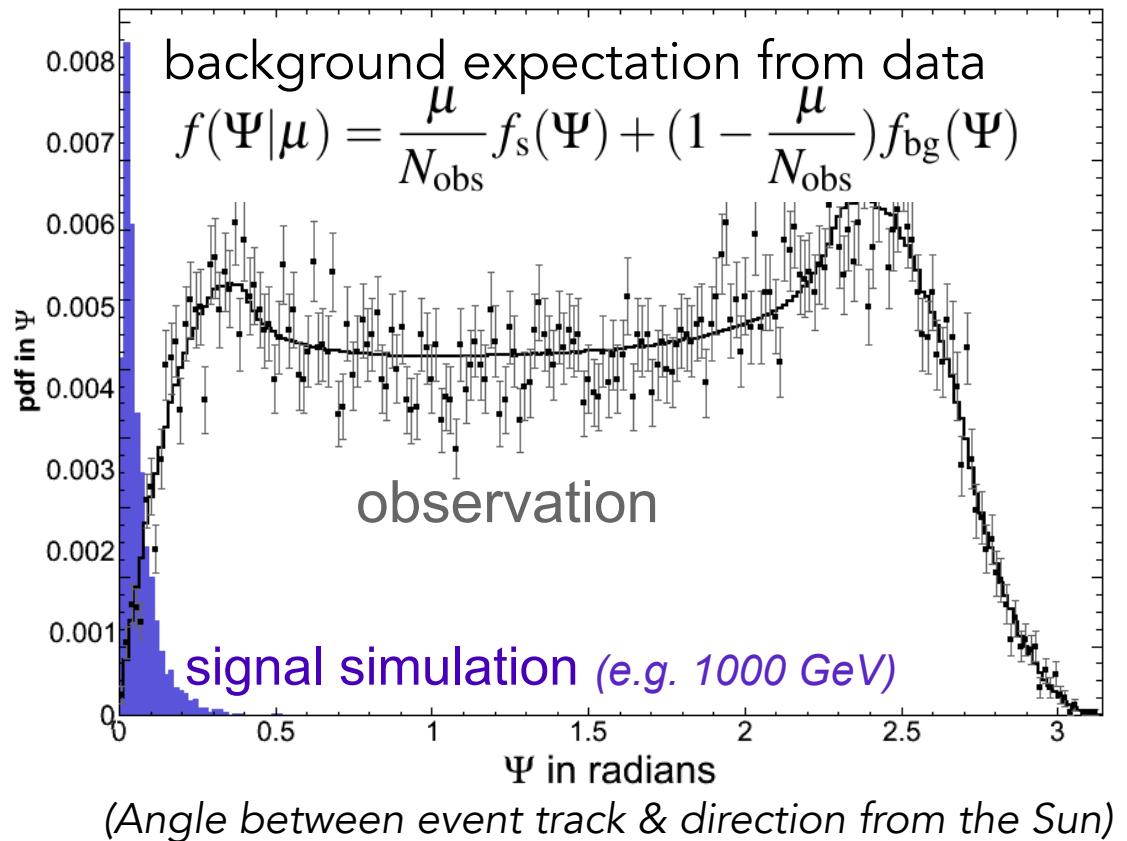


The observed angle to the Sun is fitted with *signal* and *background* pdf:s

Evaluate shape fit with log-likelihood rank (**FC**) to construct **CI** for the number of signal events μ_s

$$R(\mu) = \frac{\mathcal{L}(\mu)}{\mathcal{L}(\hat{\mu})}$$

$\mathcal{L}(\mu)$ is the pdf product over the final sample



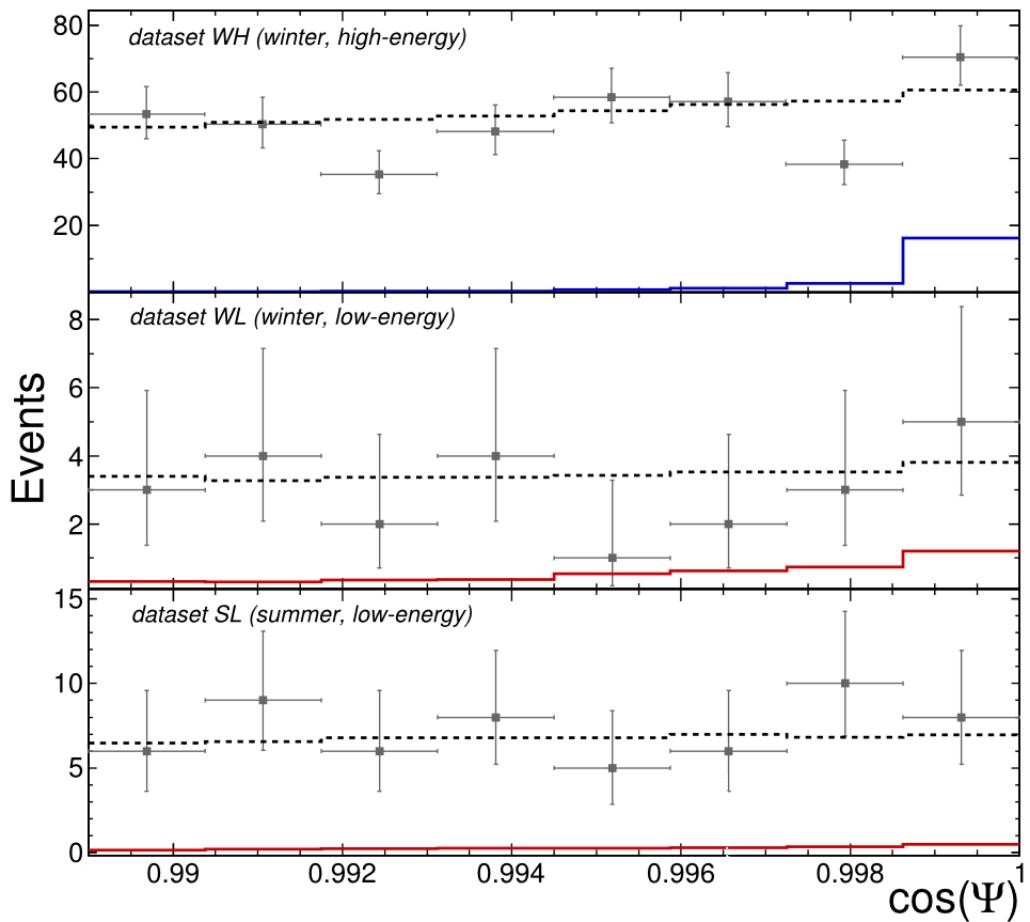
$$\mu_j = \mu \frac{T_{\text{live}}^j V_{\text{eff}}^j}{T_{\text{live}}^1 V_{\text{eff}}^1 + T_{\text{live}}^2 V_{\text{eff}}^2}$$

(scale to multiple datasets)

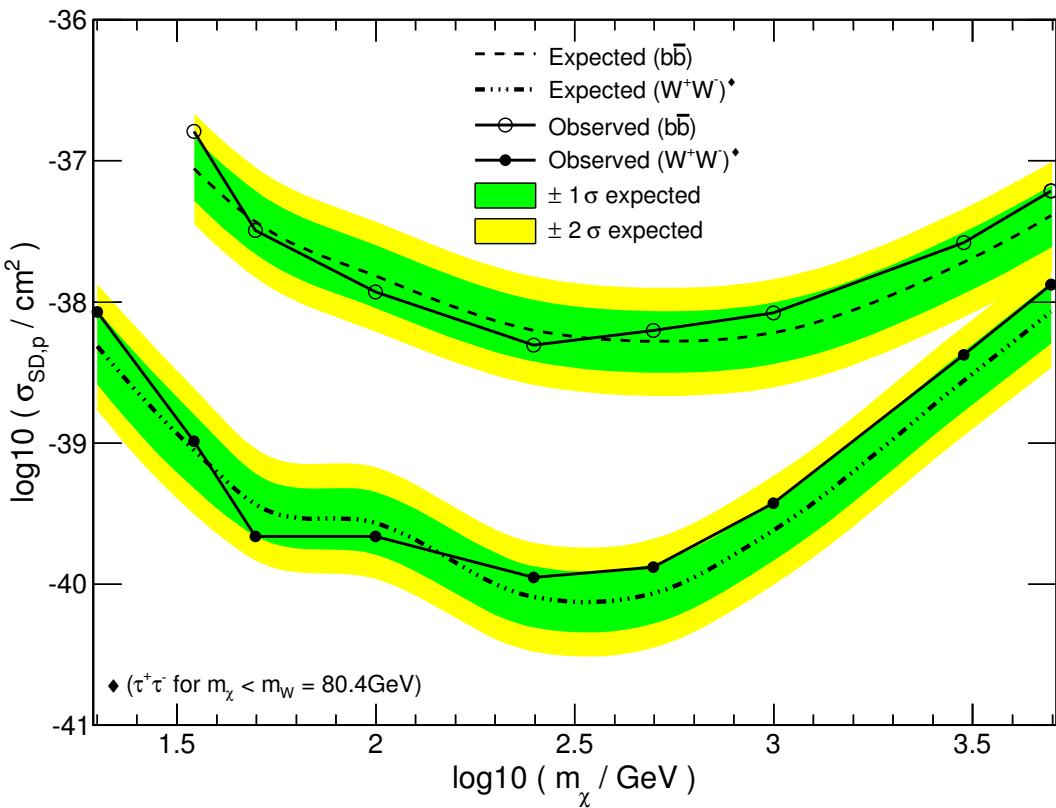
Unblinding results (observed results)



Unblinded events in different samples



Expected sens. vs. observed result

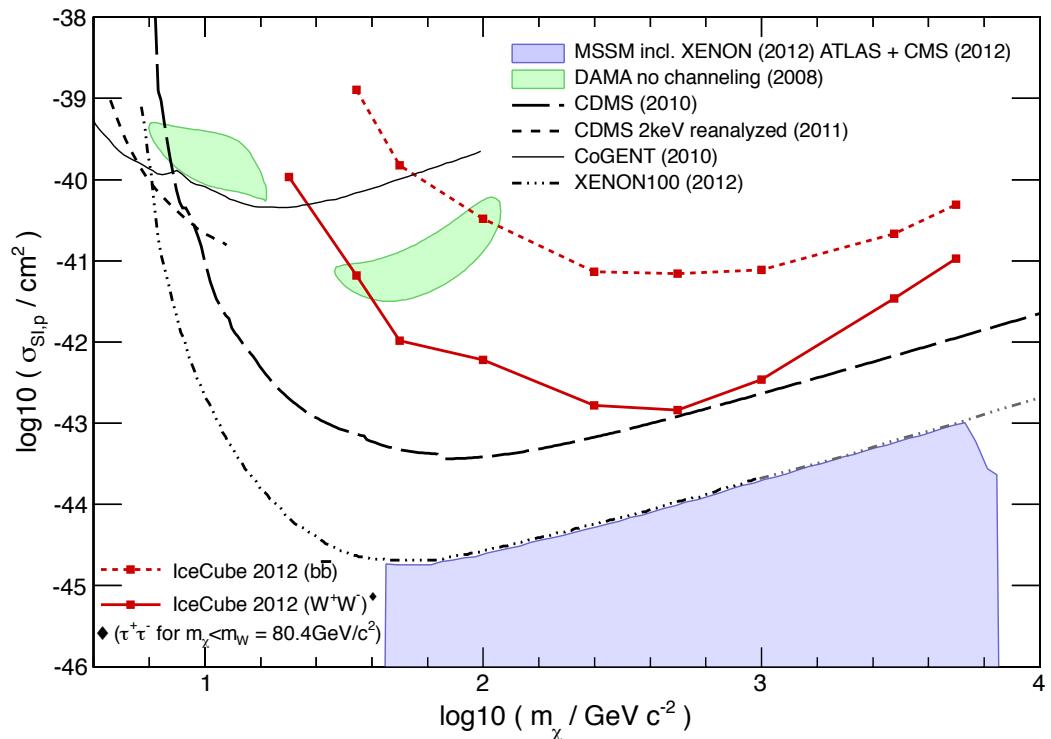


details on systematic uncertainties,
see PRL 110 (2013) 131302

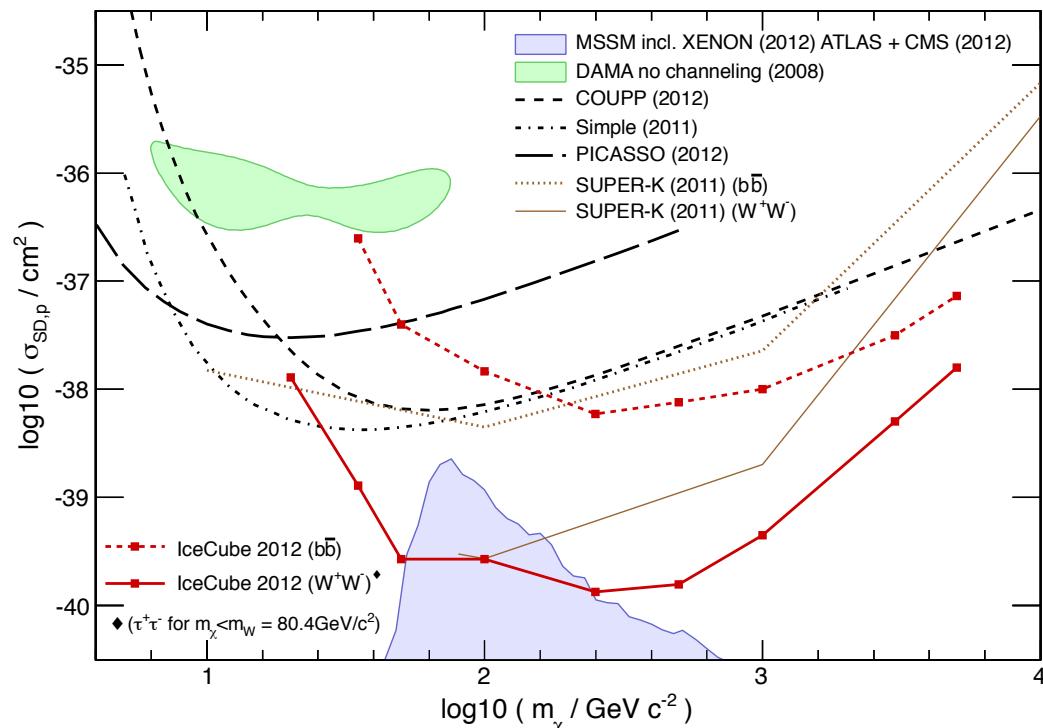
Solar WIMP search results (cross-section limit)



SI WIMP-proton cross-section limit



SD WIMP-proton cross-section limit

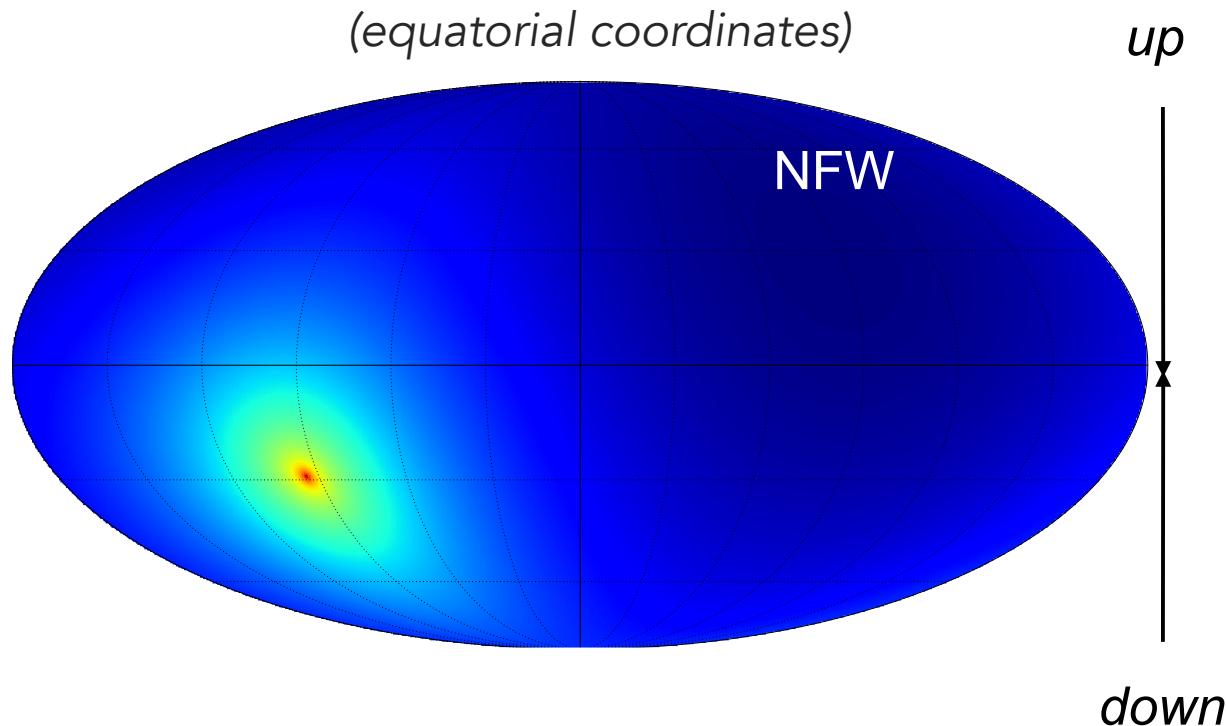


- most stringent SD cross-section limit for most models
- complementary to direct detection search efforts
- different astrophysical & nuclear form-factor uncertainties

Galactic Dark Matter searches (IceCube-79)



- Searching for neutrinos from self-annihilating *WIMPs* in the GC
- WIMP masses between *30 GeV* and *10 TeV*
- 79 string configuration of IceCube (*320 live days of 2010 - 2011*)



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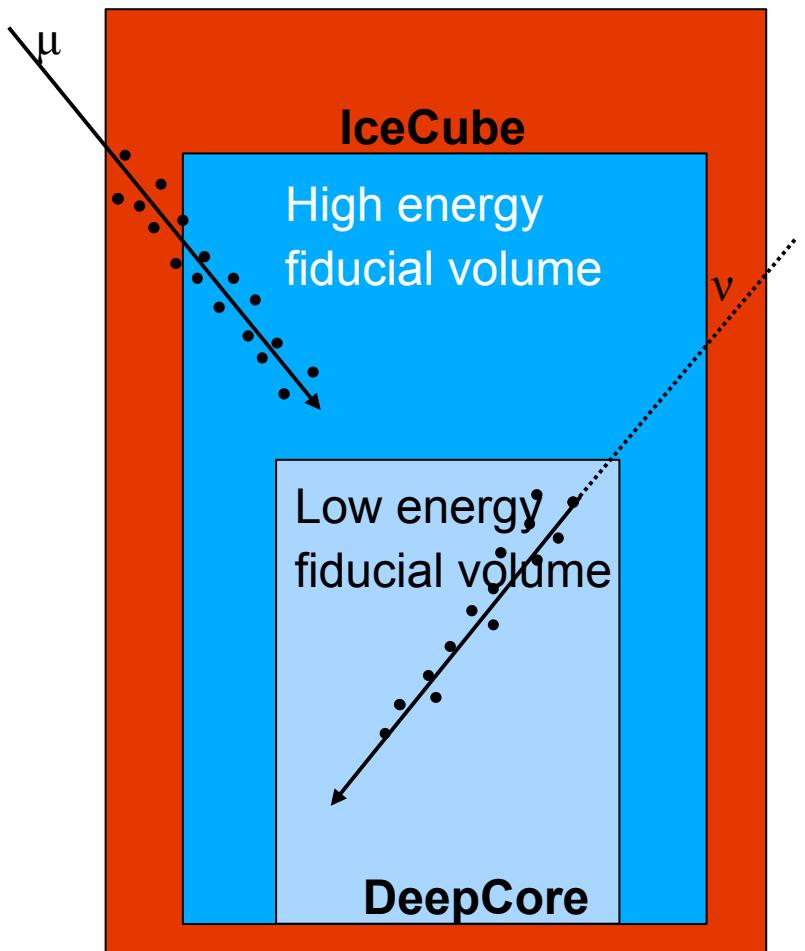
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2 independent analyses:

Low-Energy ($m_\chi < 300 \text{ GeV}$)

High-Energy ($m_\chi > 300 \text{ GeV}$)

- Identifying starting events opens up the Southern Sky for IceCube.



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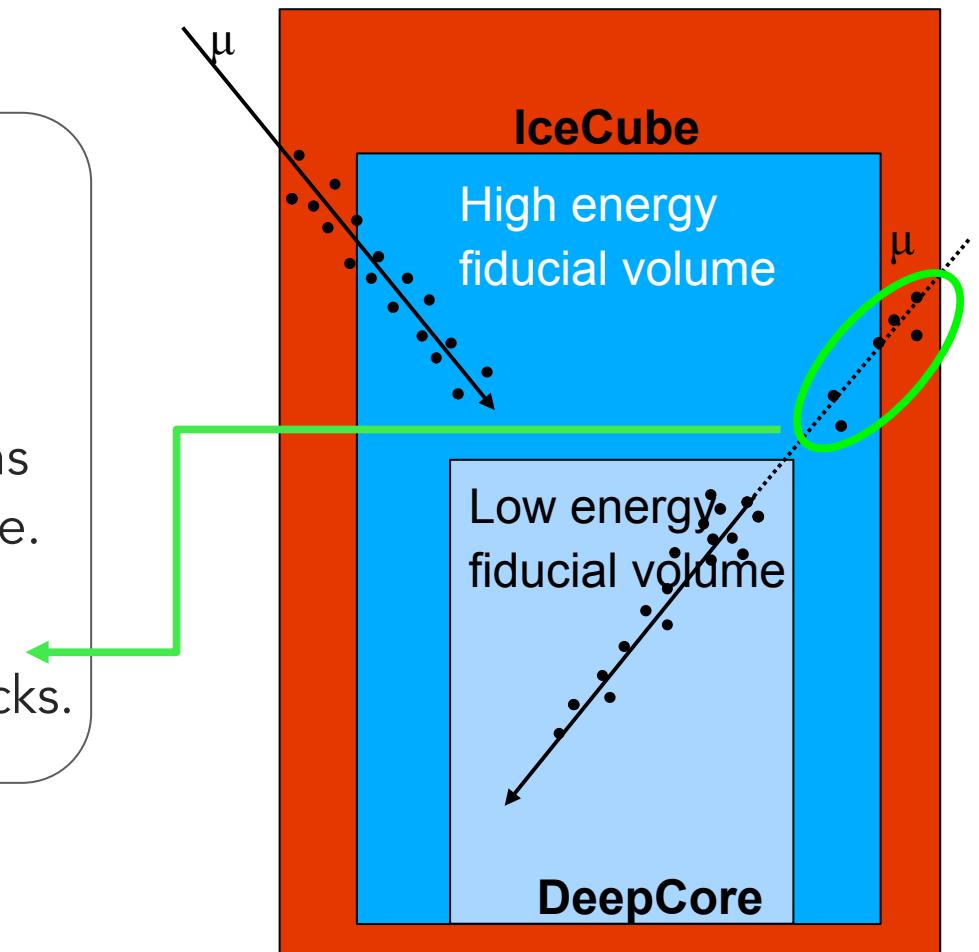
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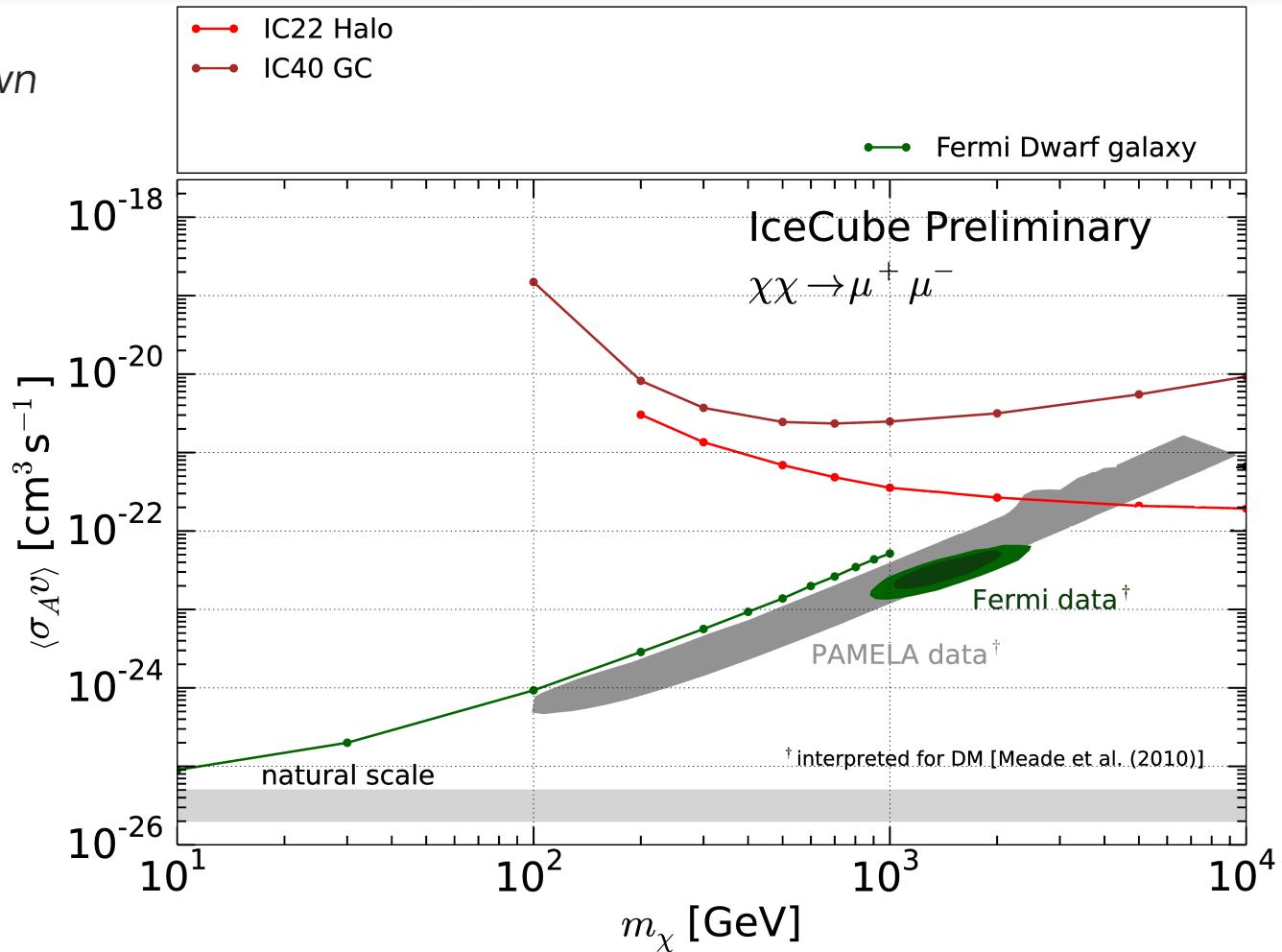
- Identifying starting events opens up the Southern Sky for IceCube.
- Both analyses rely on veto methods to reject incoming tracks.



Galactic Dark Matter searches (IceCube-79)



(IceCube results shown
for NFW profile)



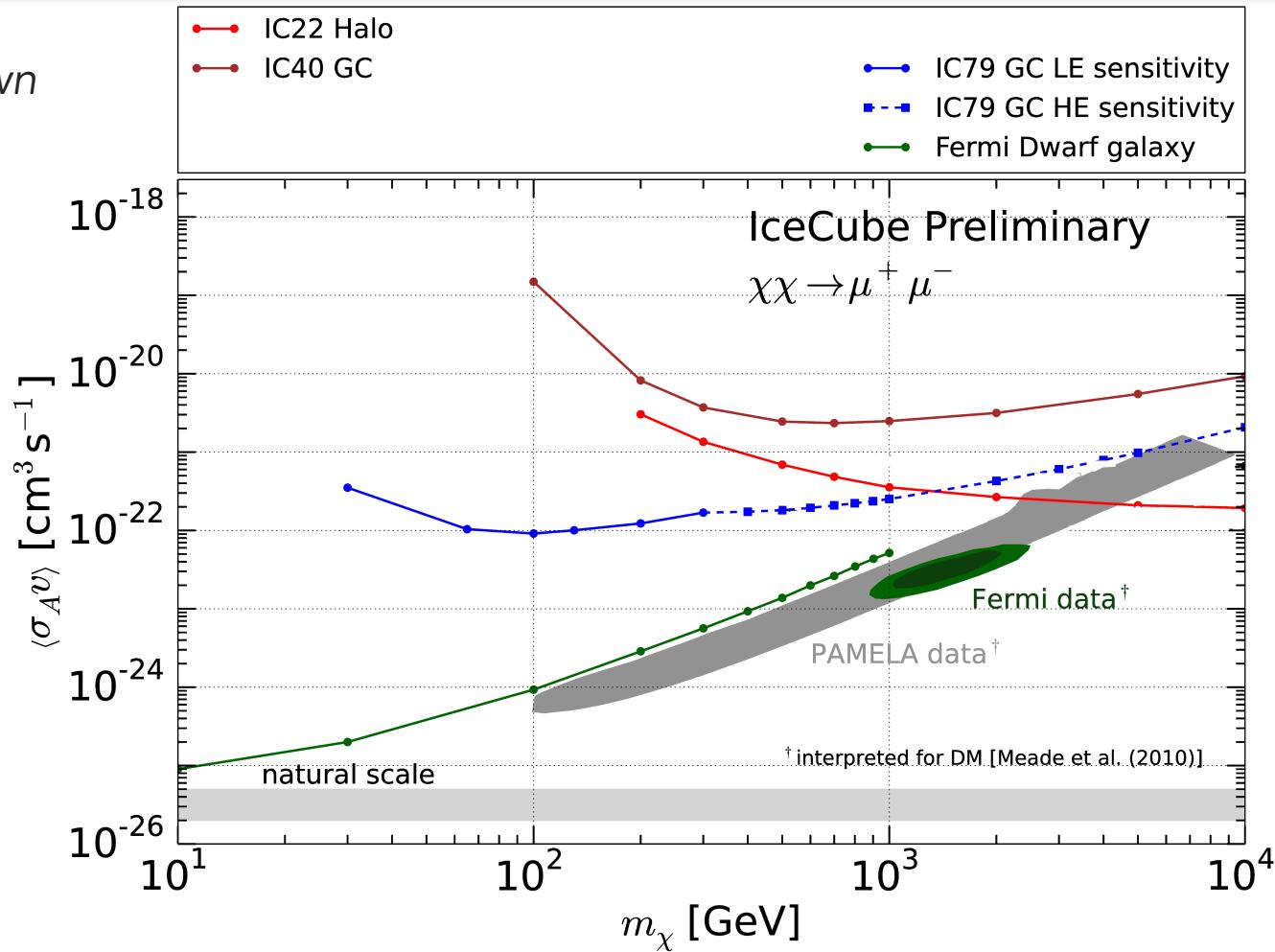
Search for many interesting
potential annihilation channels:
(Various DM-Halo models tested)

$\chi\chi \left\{ \begin{array}{l} \nu \bar{\nu}, \mu^+ \bar{\mu}, \tau^+ \bar{\tau}, W^+ W^-, b \bar{b} \\ Z^0 Z^0, Z^0 \gamma \end{array} \right.$

Galactic Dark Matter searches (IceCube-79)



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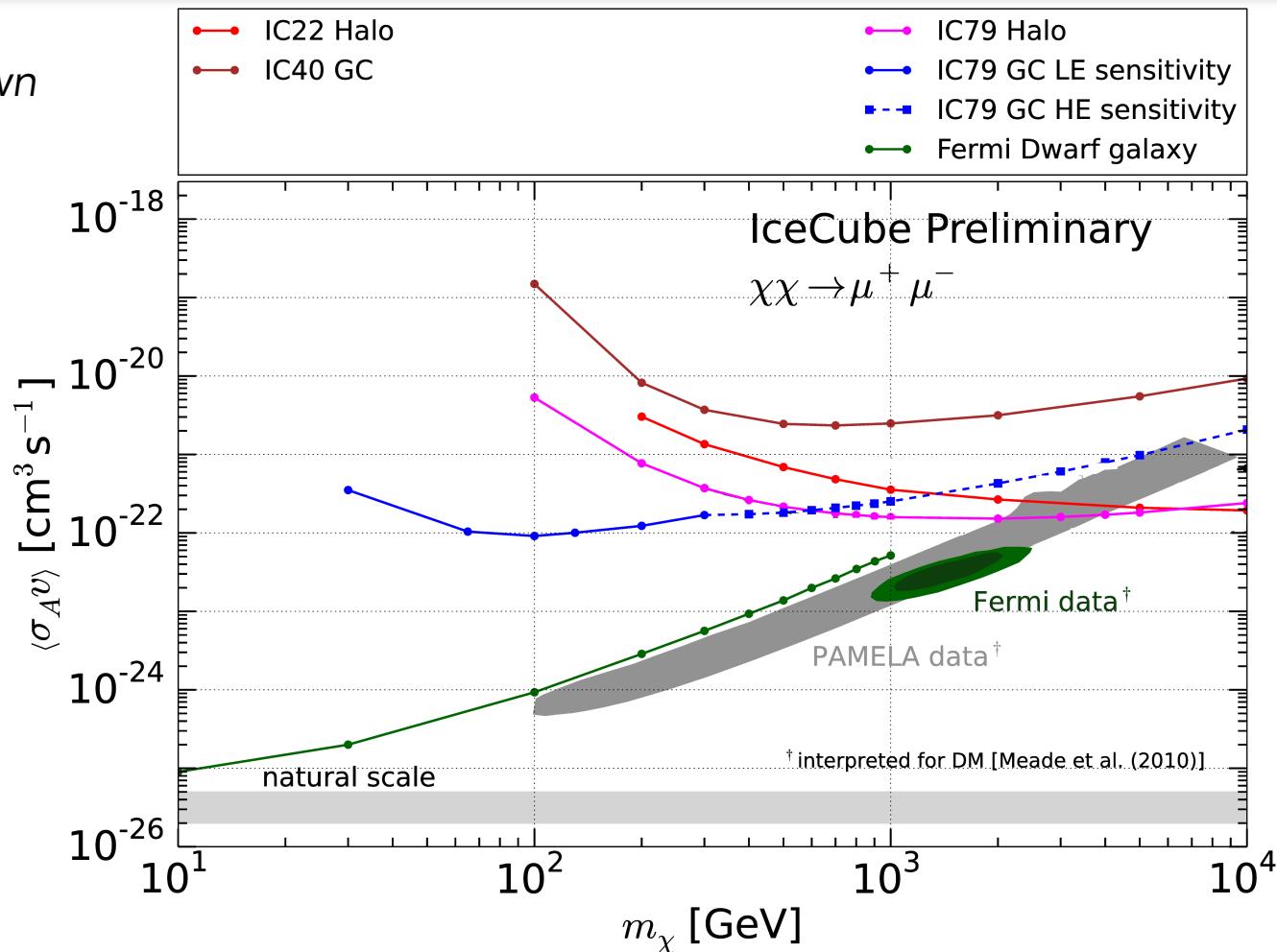
[IceCube-79 Galactic Center analysis \(sensitivity\):](#)

- First IceCube analysis looking at GC for low WIMP masses (< 100 GeV)
- Up to 4 orders of magnitude improved sensitivity @ 100 GeV
- Unblinding is going on within the collaboration

Galactic Dark Matter searches (IceCube-79)



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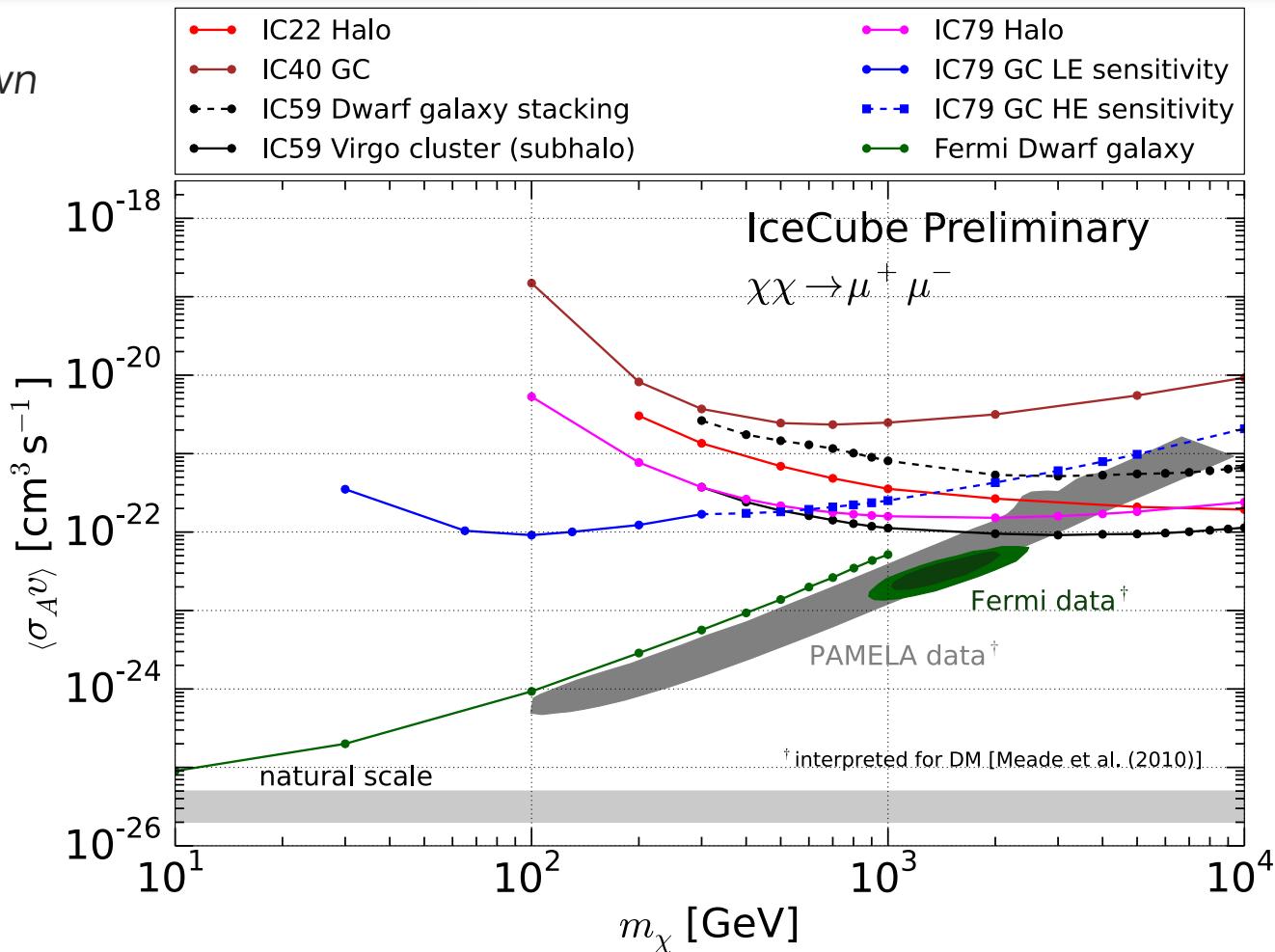
IceCube-79 Multipole analysis to search for Dark Matter in the *Galactic Halo*:

- focus on large scale anisotropies ($|l| < 100$)
- small Halo-model dependency
- results are compatible with the background-only hypothesis

Galactic Dark Matter searches (IceCube-79)



(IceCube results shown
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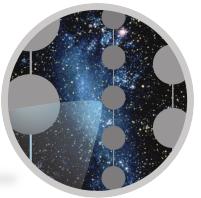


IceCube-59 Dwarf galaxy searches:

- Source stacking analysis
- Optimized size of search window

IceCube-59 Galaxy cluster analysis:

- Extended point source search
- Optimized size of search window
- Substructures taken into account



Heavy Dark Matter

- IceCube has reported 2 high-energy cascade events in 2 years of IceCube 79 + 86-string data
- consistent with νe interactions at about 1 PeV
- reported events are intriguingly close in energy

Could this be dark matter ?

- B. Feldstein, A. Kusenko, S. Matsumoto, and T. Yanagida PRD 88 (2013) 015004
- A. Arman and P. Serpico arXiv:1308.1105 (2013)

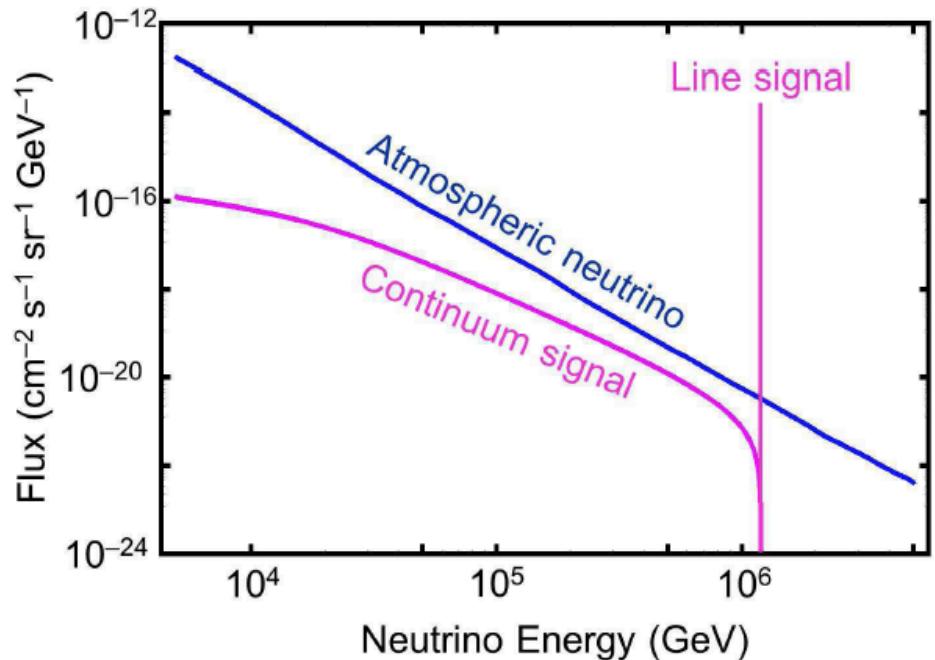
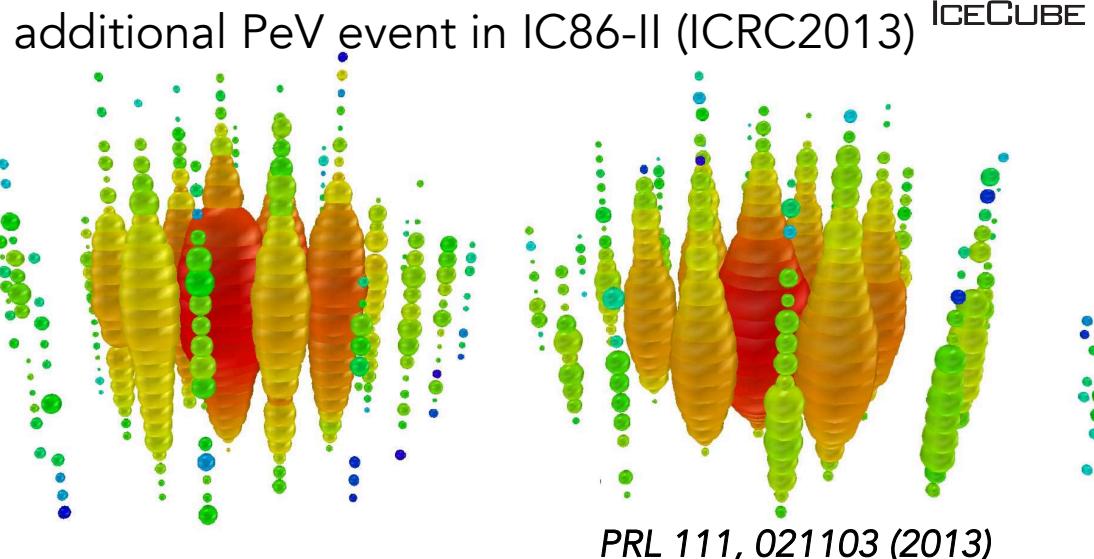
Possible evidence:

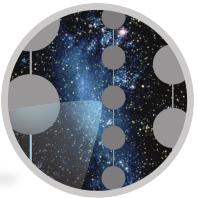
- 2.4PeV Dark Matter Particle mass
- Flux can be related to lifetime τ_{DM}

$$\tau_{DM} \simeq 1.9 N_\nu \times 10^{28} \text{ s}$$

Models:

- Singlet fermion in extra dimension
- Hidden Sector Gauge Boson
- Gravitino Dark Matter





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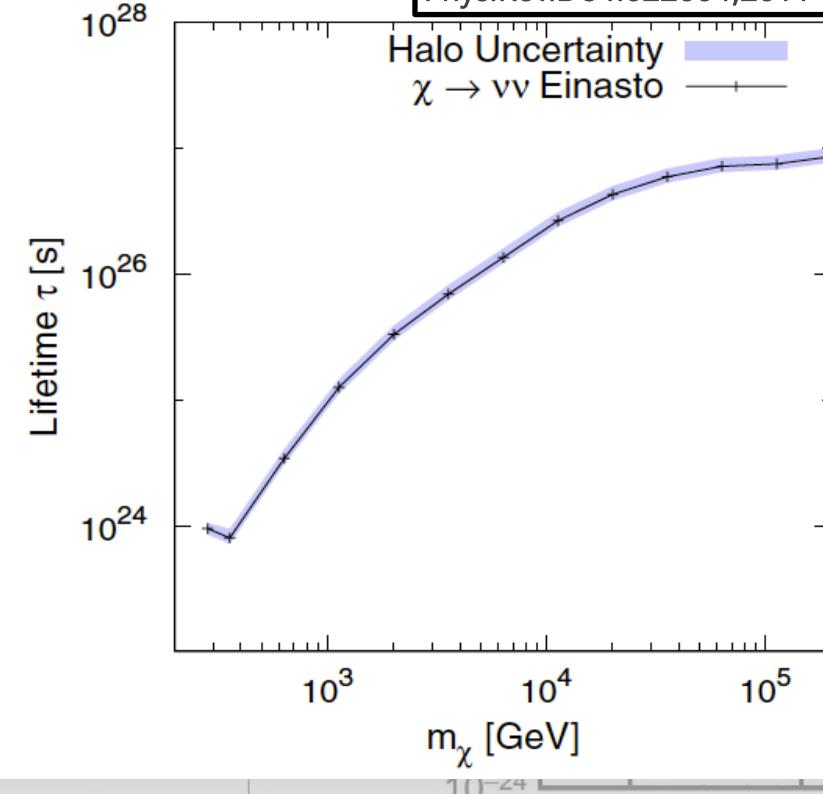
additional PeV event in IC86-II (ICRC2013)



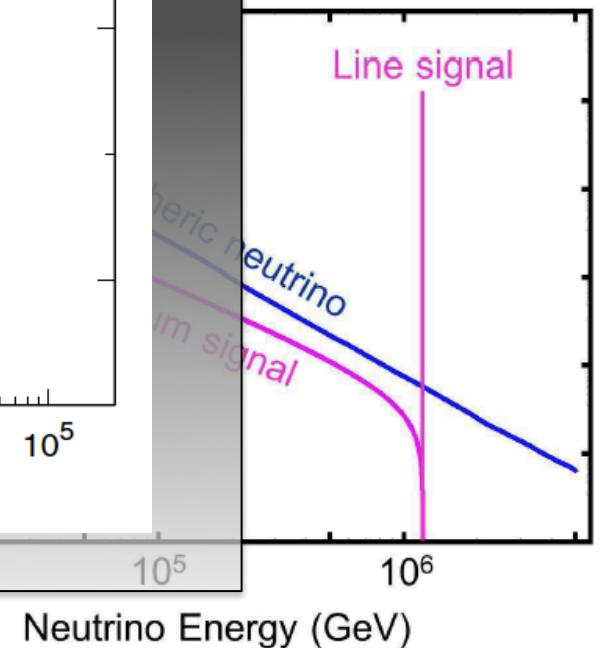
ICECUBE

IceCube bound on lifetime $\sim 10^{27}$ s

Phys. Rev. D84:022004, 2011



L 111, 021103 (2013)



Summary



- Striking WIMP signatures provide high discovery potential for indirect searches with ν , complementary to searches using other astrophysical messengers & direct detection searches
- DeepCore plays crucial role in IceCube Dark Matter analyses
- IceCube provides most stringent limits on the SD-WIMP-proton scattering cross section for most WIMP models
- Clever new ideas for detection channels and sources spur new analyses → current detectors acquire fresh data
- Future detectors with lower energy thresholds will probe region of parameter space made interesting by direct detection experiments



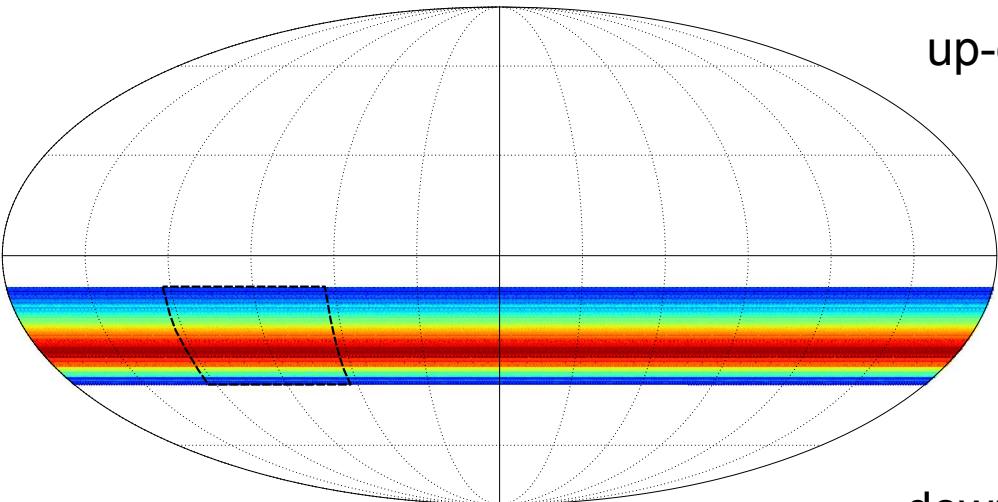
Additional slides

Galactic Dark Matter searches (IceCube-79)



2D skymap PDFs generated with healpix
(equatorial coordinates)

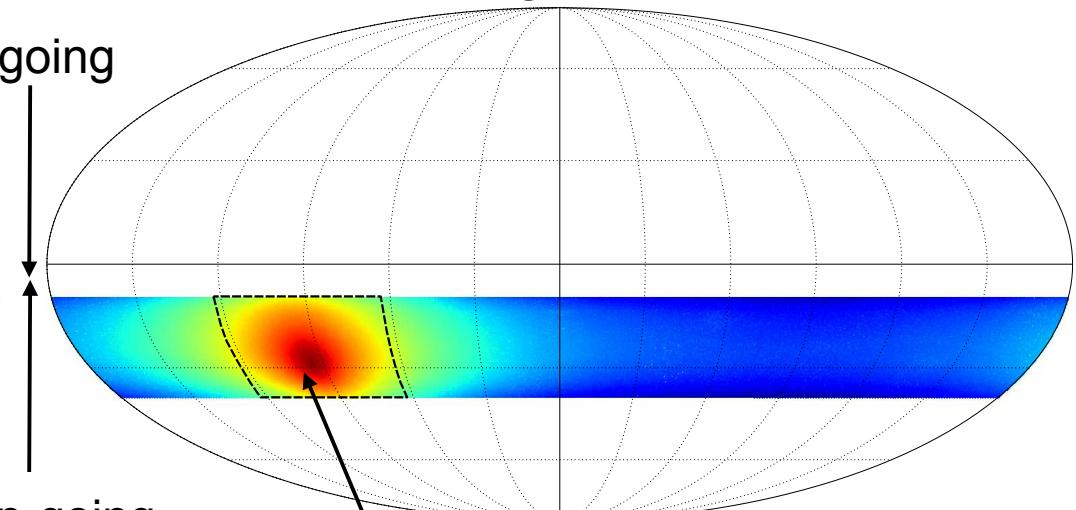
scrambled background



up-going

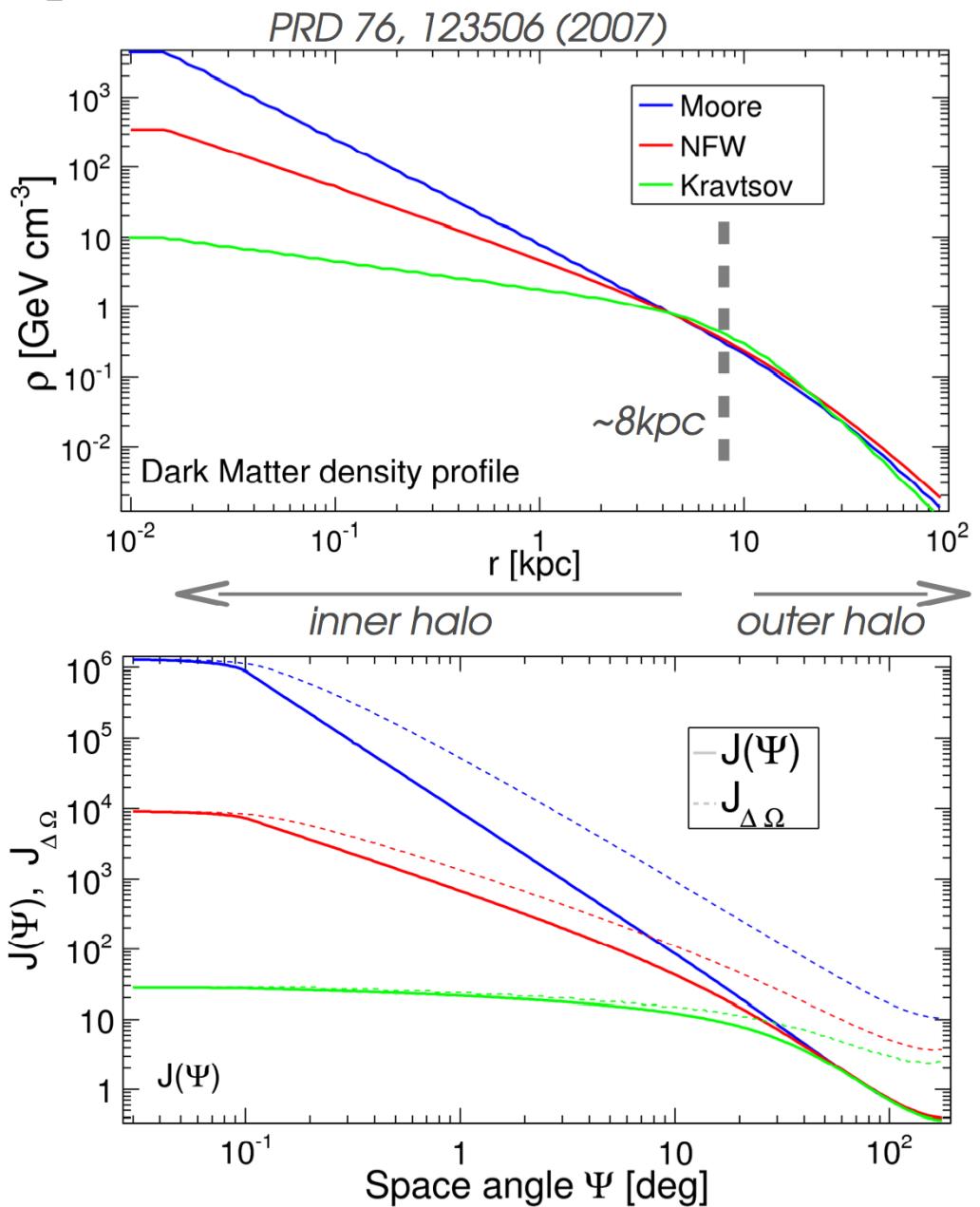
down-going

signal



Galactic Center @ 266° RA; -29° Dec

Galactic Dark Matter searches (IceCube-79)



Galactic Dark Matter searches (IceCube-79)



(IceCube results shown

for

IC22 Halo

IC40 GC

IC59 Dwarf galaxy stacking

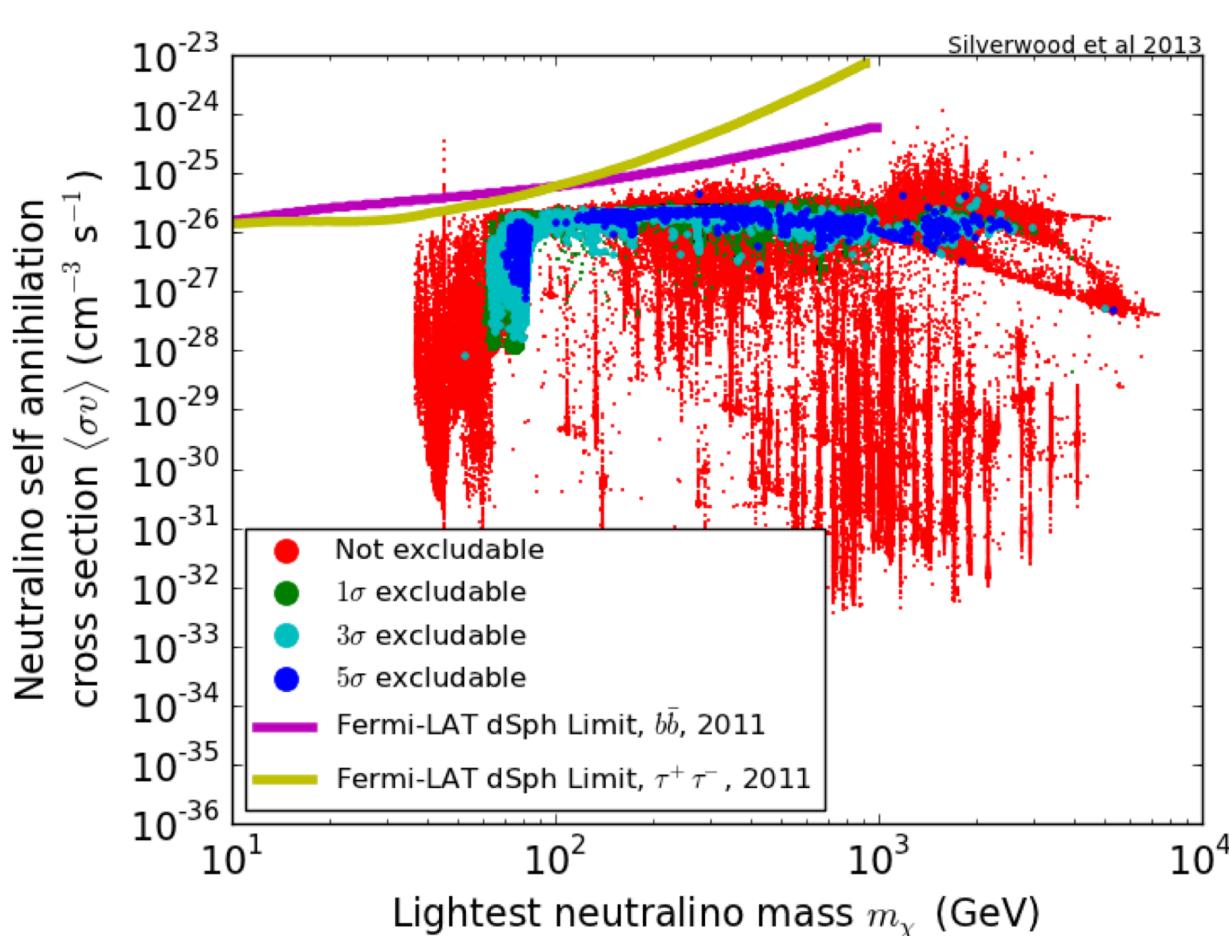
IC79 Halo

IC79 GC LE sensitivity

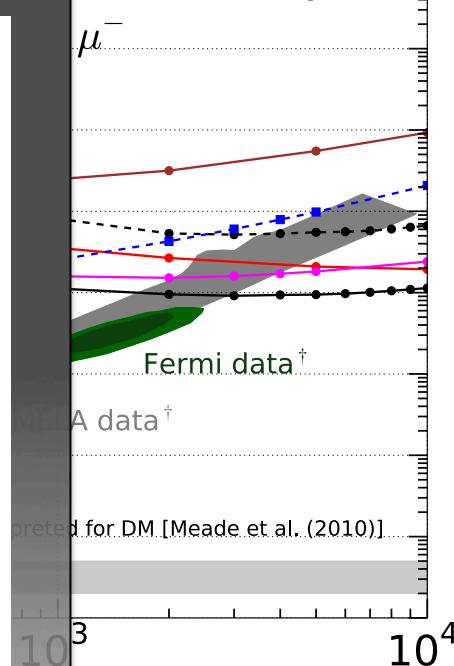
IC79 GC HE sensitivity

Fermi Dwarf galaxy

Complementarity to solar DM search

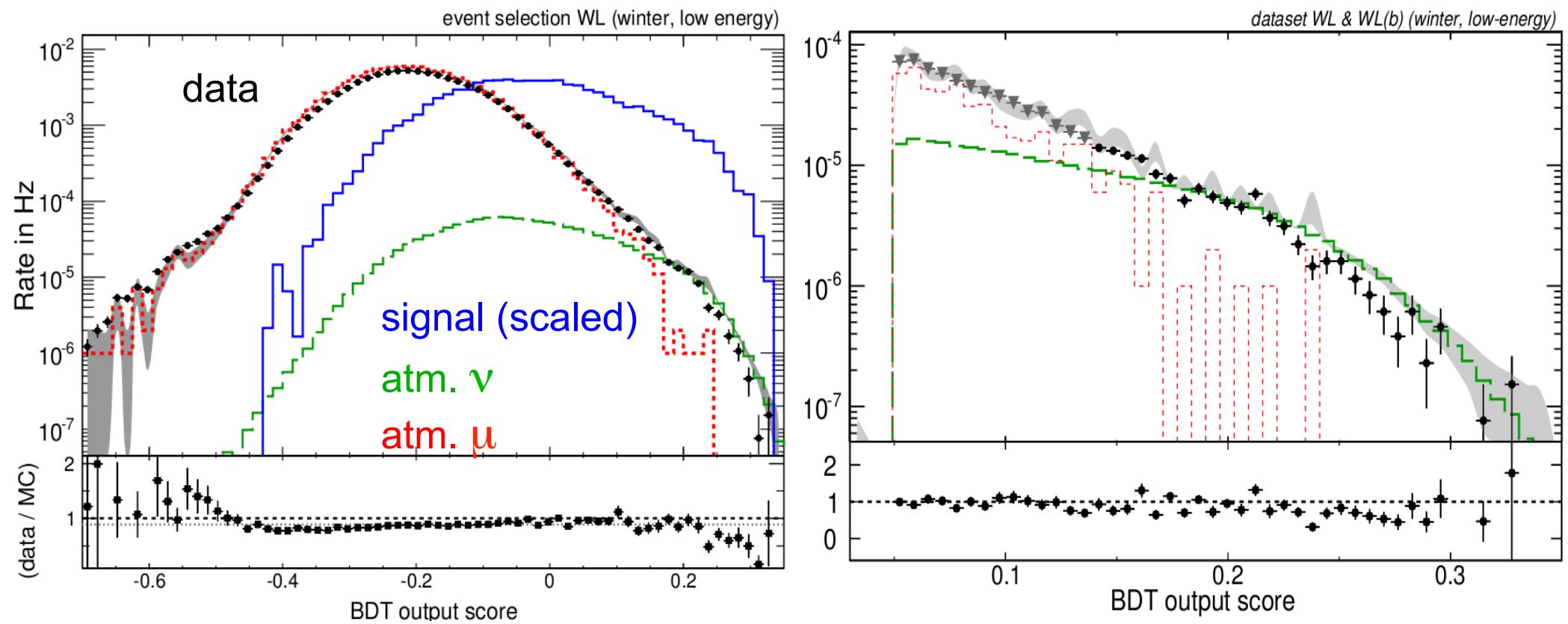


ICE CUBE Preliminary



Galaxy cluster analysis:
point source search
fixed size of search window
Substructures taken into account

Multivariate analysis step (BDT variable)

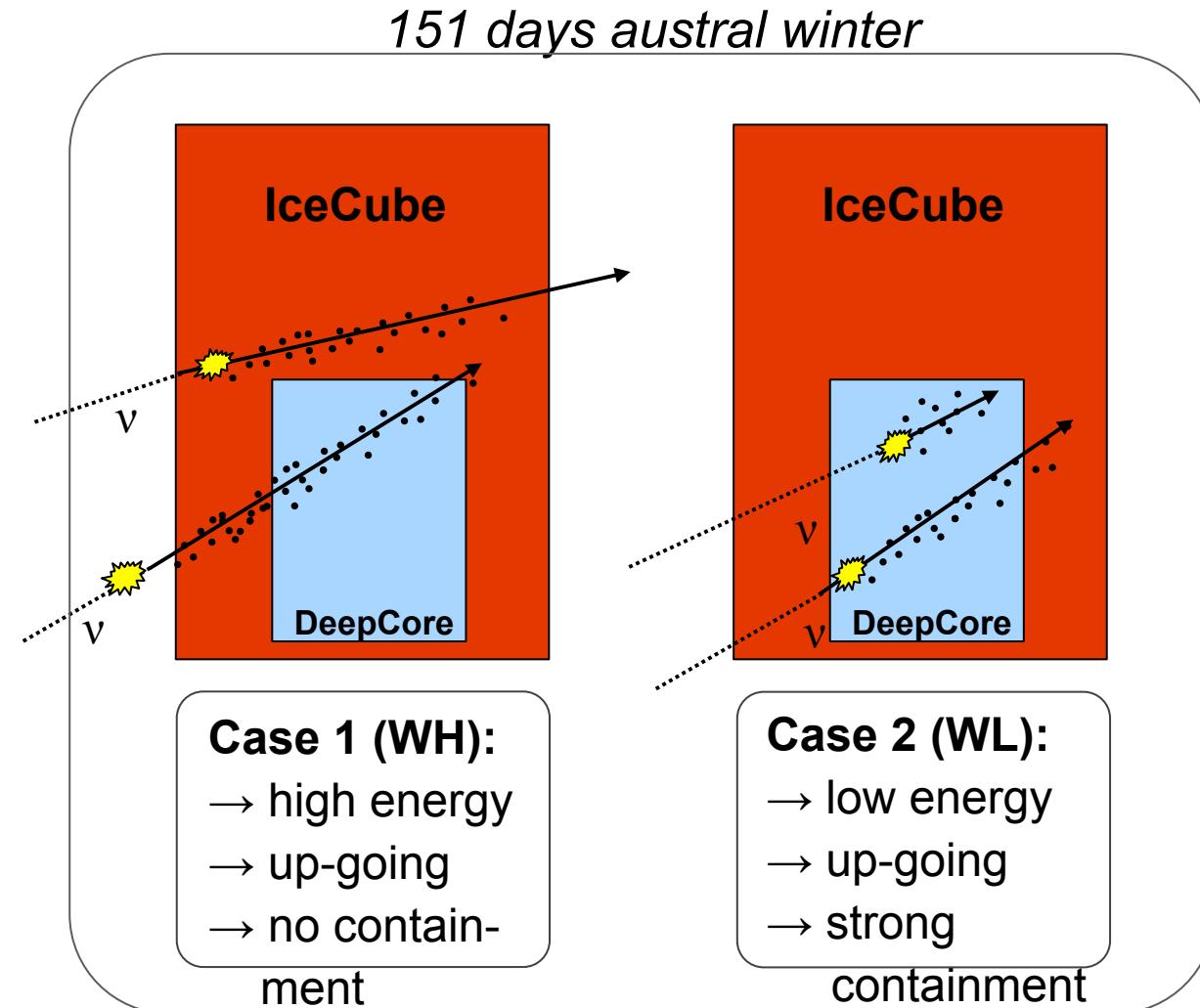


- **1 separate BDT for each event selection**
- training on off-source exp. data + separate signal simulation

IceCube-79 string analysis details



→ Different event topologies require different selection criterias to be efficient.

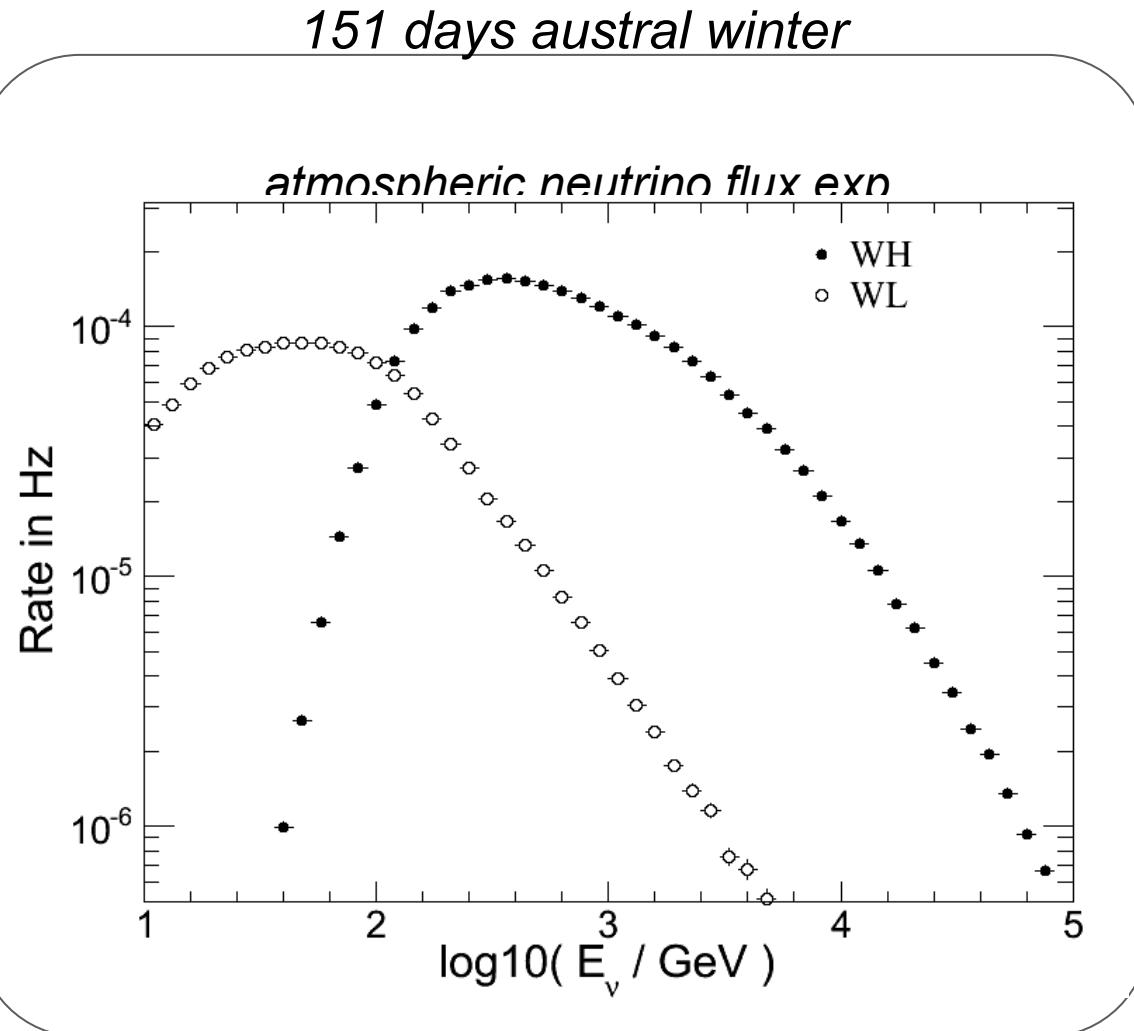


- Split data into a winter-high energy (WH) and winter-low energy (WL) event selection.
 - WL events are DeepCore dominated (more hits in DeepCore)
 - Split corrects for higher photocathode density in DeepCore

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